



THE 15TH RUSSIAN-CHINESE WORKSHOP
ON SPACE WEATHER

September 9-13, 2024, Irkutsk, Russia

NSSC
中国科学院国家空间科学中心
National Space Science Center . CAS



MUSER observation and joint study with SRH



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September 11, 2024, Irkutsk,
Russia

Introduction of MingAnTu Station



under the Meridian-II Project



1. Interplanetary Scintillation telescope
2. MUSER-L, Meter-Decameter Spectral Radioheliograph
3. MUSER-I, Decimeter Spectral Radioheliograph
4. MUSER-H, Centimeter Spectral Radioheliograph (to be upgraded)
5. Centimeter Spectragraph , 6. Decimeter Spectragraph

1. Interplanetary Scintillation telescope

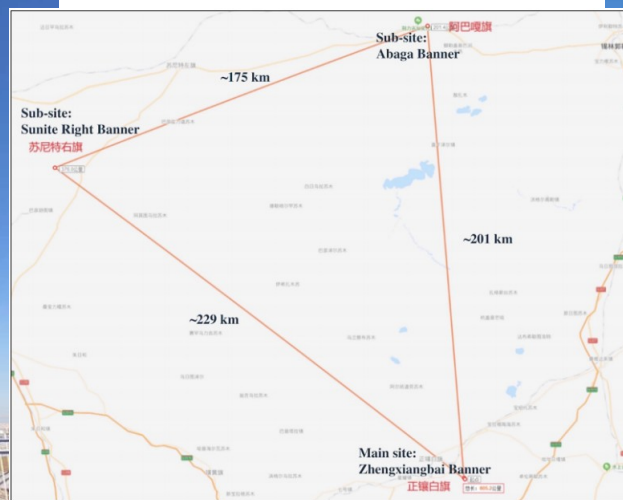
Main performance of the Chinese IPS telescope.

Antennas: Cylinder: 3 *140m*40m, parabolic: 2×U30 m

frequency: 327 MHz, 654 MHz, 1420 MHz (only for 30 m dish antennas)



30 meter dish antennas



MUSER-I

Frequency : **0.4-2.0 GHz**

Antenna : Φ 4.5m \times 40

FOV: $>2.33^\circ$ (2 GHz)

Channel: 64

Freq. resolution : 25 MHz

Cadence : 25 ms

Spatial resolution : 10.3"-51.6"

Image dynamic range : >25 dB

polarization : R, L

MUSER-H

Frequency: **2.0-15.0 GHz**

Antenna : Φ 2.0m \times 60

FOV: $>0.7^\circ$ (15 GHz)

Channel: 528

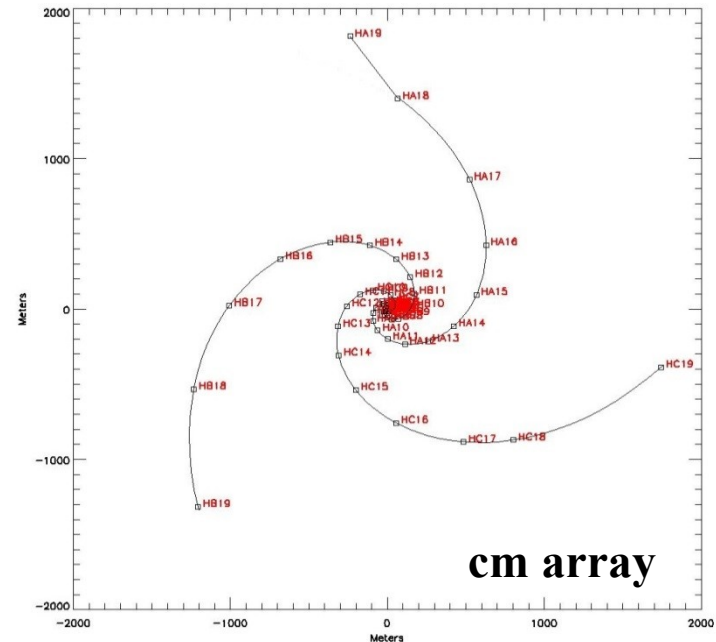
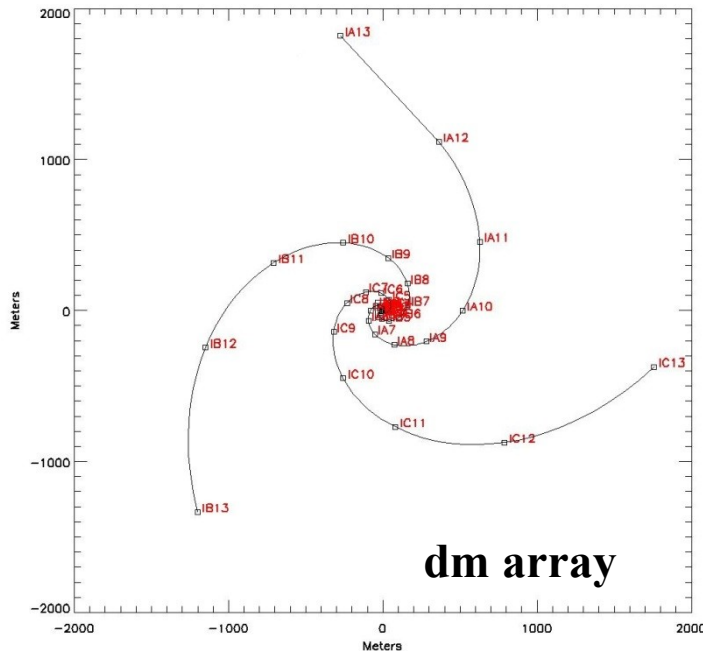
Freq. resolution : \sim 25 MHz

Cadence : \sim 200 ms

Spatial resolution : 1.4"-10.3"

Image dynamic range : >25 dB

polarization : R, L



Max Baseline: **3 km**, aperture-synthesis technology

MUSER-L

Frequency : **30-400 MHz**

Antenna : $(100 + 124) \times \text{LPDA}$

FOV: $\sim 120^\circ$

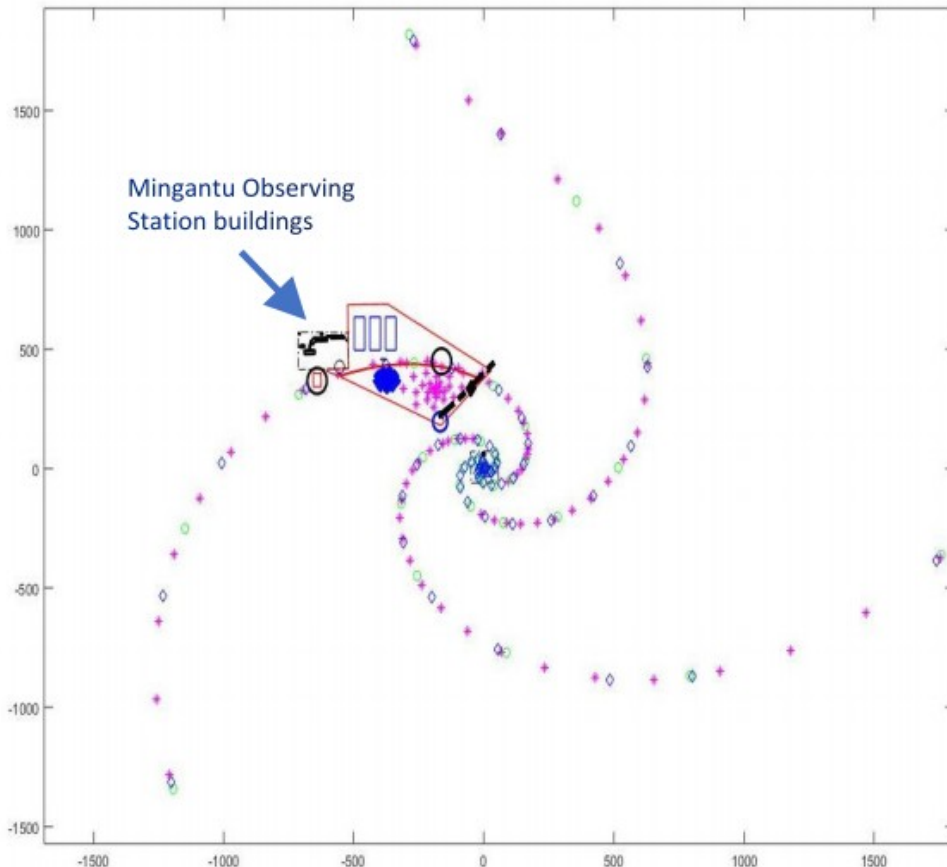
Channel: 384

Cadence : ~ 100 ms

Spatial resolution : $14'.0 - 1'.0$

Image dynamic range : $>25\text{dB}$

Polarization : Dual linear H, V

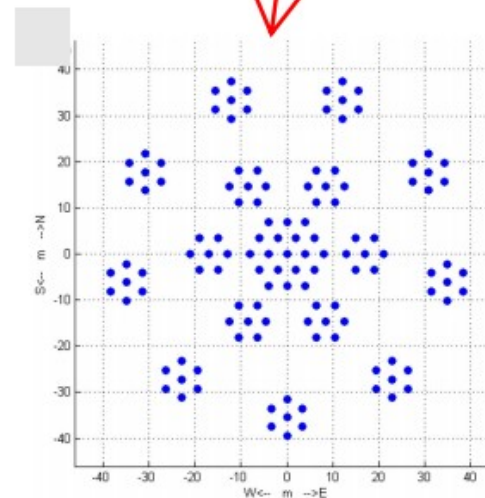
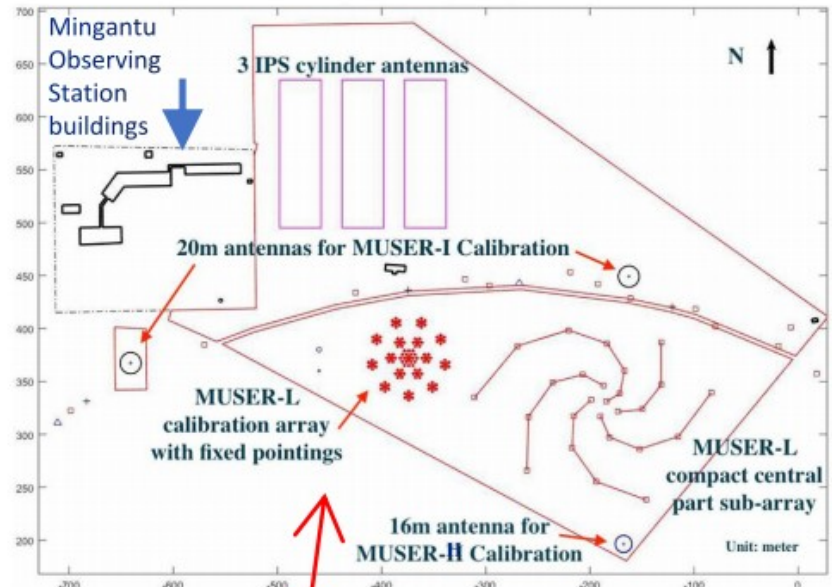


blue & green

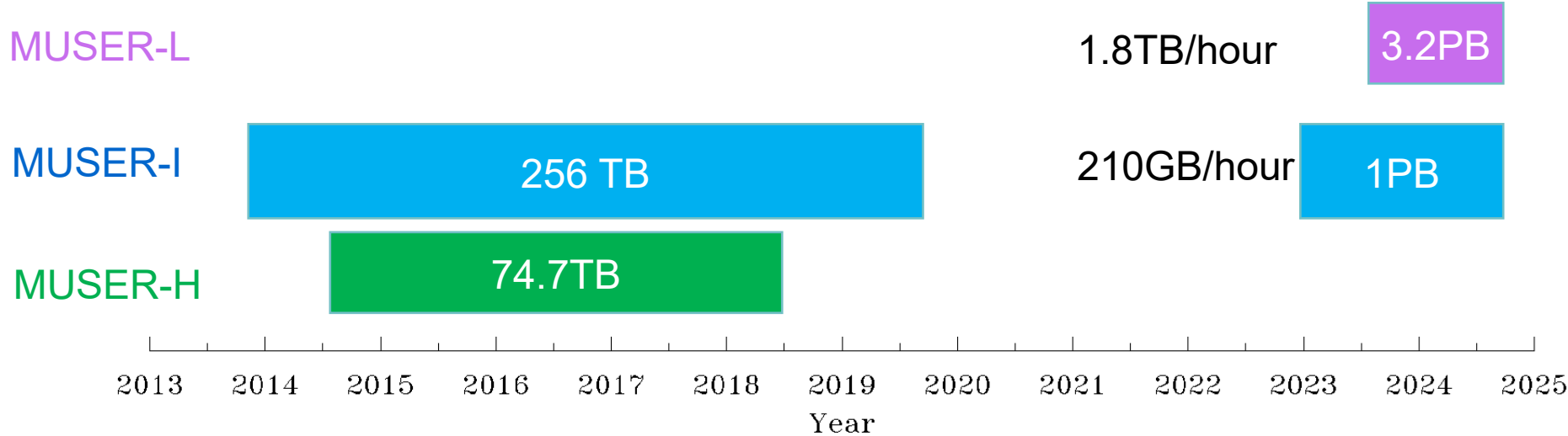
----- MUSER-I/H

pink:

-----MUSER-L



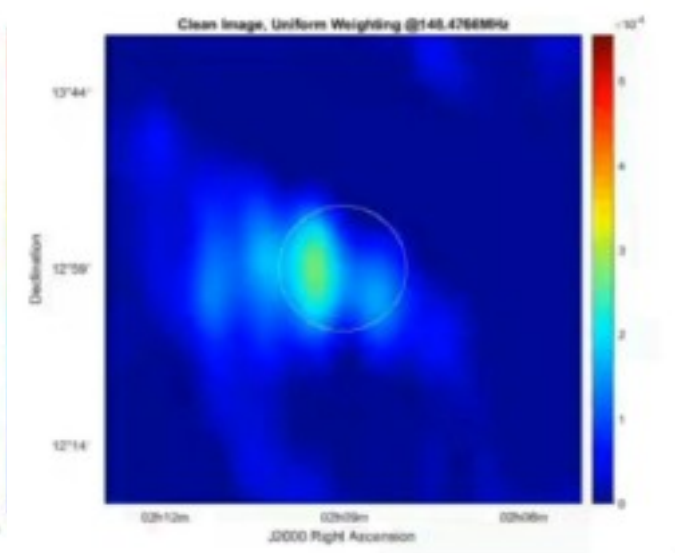
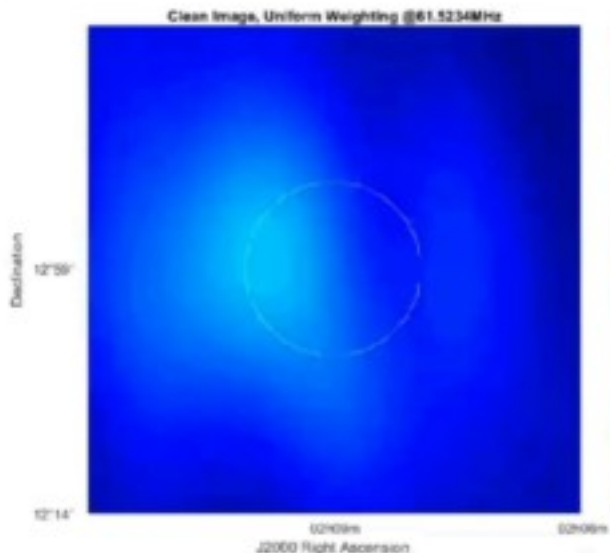
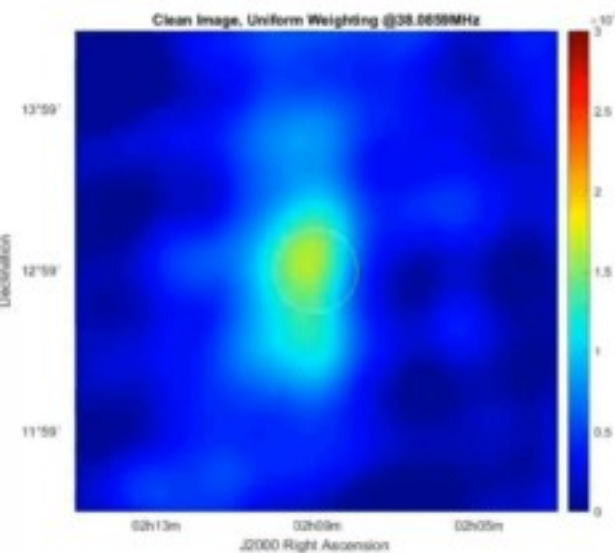
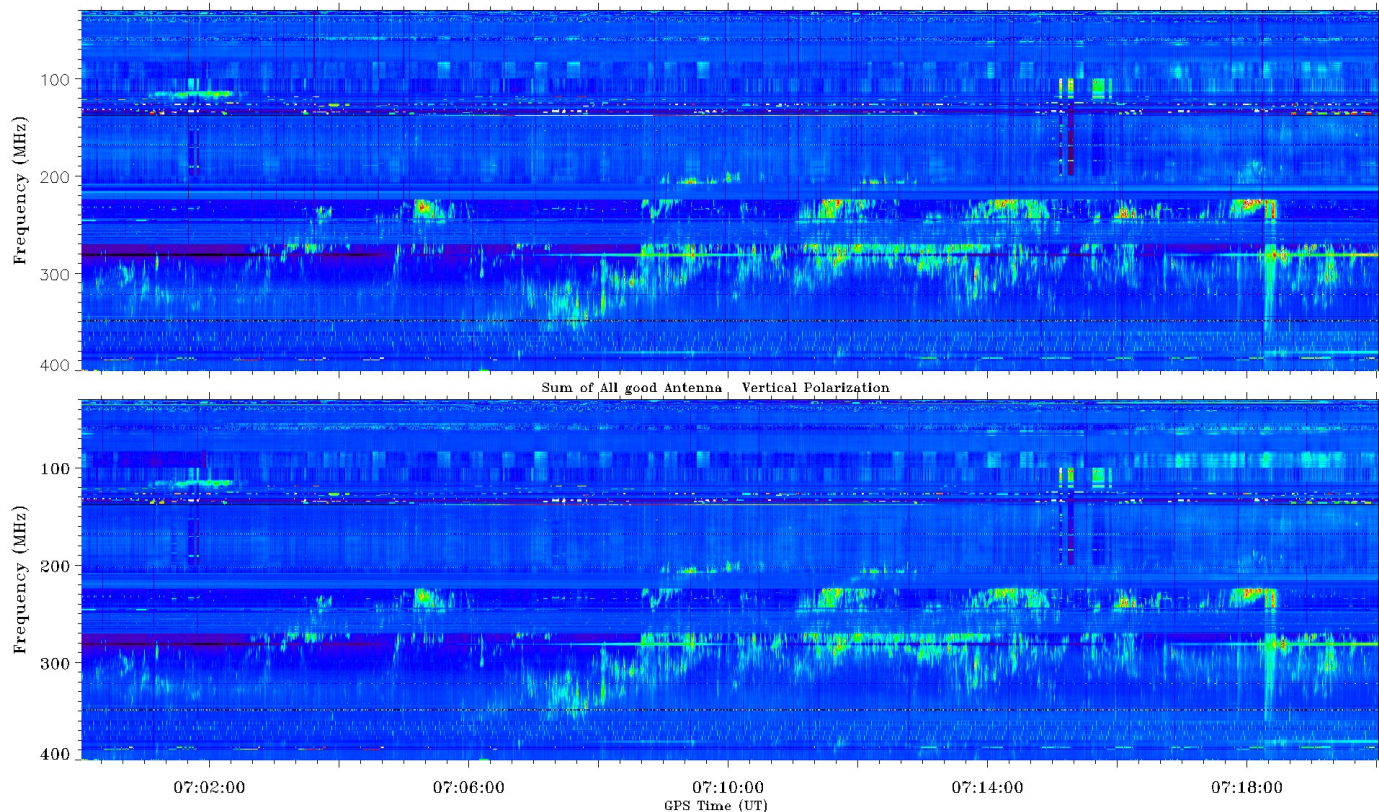
MUSER Observation



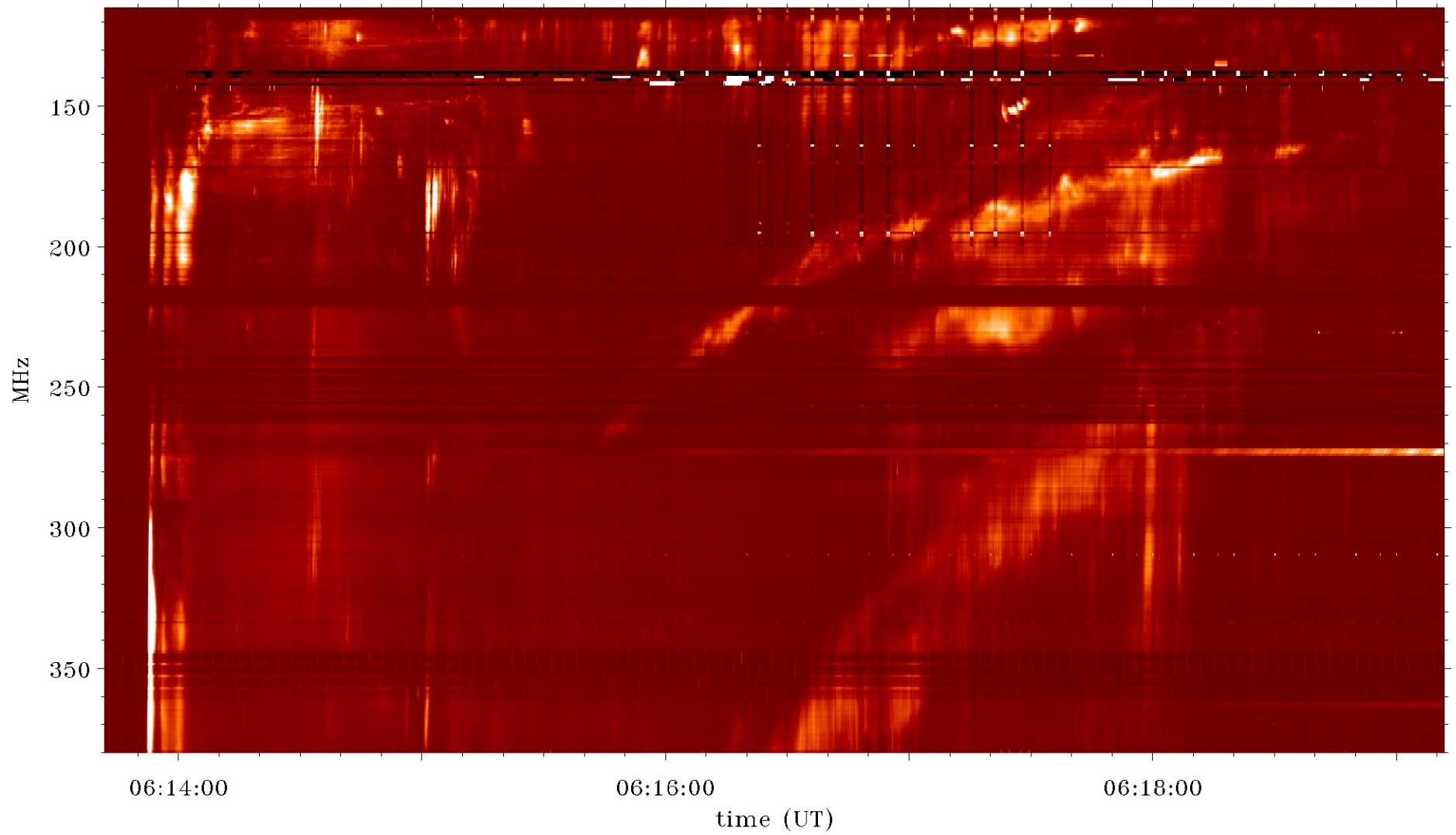
MUSER - I

frequencyband	0.4 - 2.0 GHz	0.4 - 2.0 GHz	0.4 - 2.0 GHz	1.6 - 2.0 GHz
frequency res.	16MHz	4MHz	1MHz	16MHz
time cadence	25ms	25ms	25ms	3.125ms
polarization	L, R	L, R	L, R	L
Data	2TB/10h	8TB/10h	16TB/10h	2TB/10h

MUSER-L observation



MUSER-L Spectrum on 2024-05-04



MUSER-I observation (before 2019)

94 bursts

2014-10-22, 05:13-05:16UT, M2.7 MUSER2

2014-11-11, 04:20-04:34UT, C3.4 flare and a limb eruptions

2014-12-17, 04:25-04:45UT, M8.7, QPP and radio spike burst groups

2015-06-20, 01:24-02:29UT

2015-06-21, 01:17-03:43UT

2015-11-01, 03:35-03:44UT

2015-11-22, 04:51-04:52, up and down drift burst with C1.2

2016-07-23, MUSER2

2016-08-29, 03:44-03:48UT, C3.0

2017-02-28, 06:29-03-06:29:07, B-flare

2017-04-03, 00:45-01:30, 2 groups of QPP in the rising phase of M1.2 flare

2017-05-31, 10:02-10:06, B7.0

2017-06-02, 01:31-01:50, C-class

2017-06-03, 09:37-09:38, C-class

2017-06-08, 05:46-05:47, B4.0

2017-07-03, B9.0, to be done, weak

2017-07-09, 03:40-03:43 M flare and several C-class flares

2017-07-10, 03:50-05:50, B6.0, 09:48-09:55, B3.0

2017-07-13, 05:19-05:20, B5.0

2017-07-14, 01:20-04:00 M-flare

2017-07-16, 01:20-04:00 C-flare

2017-07-17, 07:24-07:26, C2.0

2017-07-31, B9., to be identify

2017-08-01, C-flare

2017-08-15, 04:30-

2017-08-16, 08:27-

2017-08-20, 01:45-

2017-08-21, 02:30-

2017-08-22, 04:00-

2017-08-23, C1.4,

2017-08-24, 04:29-

2017-09-04, 01:11-

2017-09-05, 03:40-

2017-09-06, 07:30-

2017-09-07, 02:00

2017-09-08, 01:38

2017-09-09, 05:15-

2017-09-10, 06:30-

2017-09-12, 07:2

2017-09-26, 02:3

2018-03-30, 07:58-

2019-03-08, 03:00-

2019-05-05, 06:40-

2019-05-06, 05:00-

2019-05-07, 01:38-

2019-05-09, 05:40-

284-6

M. H. Zhang et al.: Observational Results of MUSER during 2014-2019

Table 2 Observational Results of Solar Radio Bursts

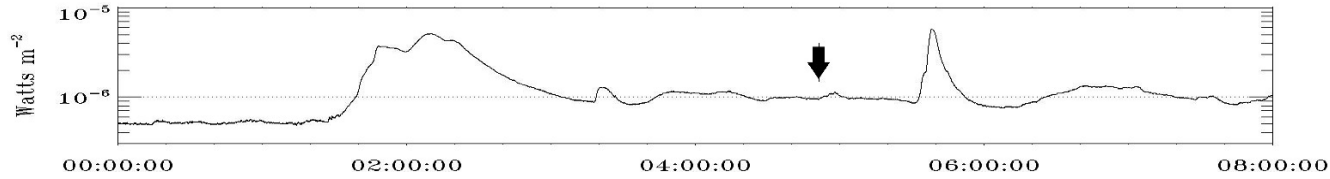
Event	Start	Stop	III-like	QPP	IV	Drift	Spl	Lac	Fib	Um	Class/level
2014 Nov 11	3:59(3.28)	4:00(4.32)	✓	✓							C1.1
	4:16(4.11)	5:05(5.26)	✓	✓							C3.4
	5:32(5.02)	5:53(6.04)	✓	✓							C2.1
	6:13(6.02)	6:30(7.14)	✓	✓							C3.1
	7:59(7.27)	8:00(8.41)	✓	✓							C2.3
2014 Dec 17	1:46(1.11)	2:27(2.27)	✓	✓	✓	✓	✓	✓	✓		M1.1
	4:22(3.55)	5:22(5.59)	✓	✓							M8.7
	5:54(4.52)	5:57(6.23)	✓	✓							B6.3
	1:13(0.51)	1:22(2.00)	✓	✓							B6.9
2015 Jun 20	2:08(2.00)	2:27(2.39)	✓	✓							C2.1
2015 Jun 21	1:17(0.32)	6:13(3.32)	✓	✓	✓	✓	✓	✓	✓		M2.0 M2.6
2015 Nov 4	3:57(3.25)	4:04(3.33)	✓	✓							C1.1
	4:52(4.20)	4:56(5.29)	✓	✓							C1.0
	4:51(4.23)	4:55(5.30)	✓	✓							C1.1
2015 Nov 22	5:35(5.01)	5:38(6.11)	✓	✓							C5.6
2016 Aug 29	3:44(3.04)	3:50(4.32)	✓	✓							C2.2
2016 Sep 22	3:26(2.33)	3:27(4.02)	✓	✓							B4.7
2016 Nov 29	7:07(6.33)	7:09(7.42)	✓	✓							C7.5
2017 Feb 28	6:28(5.53)	6:30(7.02)	✓	✓							B3.5
2017 May 1	3:58(3.25)	4:00(4.35)	✓	✓							B9.9
2017 Jun 2	1:31(0.59)	1:48(2.21)	✓	✓							B8.5 C1.1
	9:49(9.17)	9:51(10.26)	✓	✓							B16.6
2017 Jun 3	4:08	4:10	✓	✓							level-B
	9:37(9.07)	9:55(10.35)	✓	✓							C2.1
2017 Jun 22	8:24	8:26	✓	✓							level-A
2017 Jul 10	1:23	1:26	✓	✓							level-B
	3:51(3.39)	5:06(7.20)	✓	✓							B5.4
2017 Jul 12	9:48	9:53	✓	✓							level-B
	8:18	8:25	✓	✓							level-B
2017 Jul 14	1:18(0.37)	2:03(5.54)	✓	✓							M2.4
2017 Jul 16	1:58(1.28)	2:13(2.27)	✓	✓							C3.1
2017 Jul 17	7:24(7.13)	7:25(8.16)	✓	✓							B3.3
2017 Jul 18	4:30(3.54)	4:33(5.03)	✓	✓							B8.7
2017 Aug 15	5:05(4.30)	5:07(5.39)	✓	✓							B7.1
2017 Aug 16	8:27(7.59)	8:34(9.10)	✓	✓							B5.2
2017 Aug 20	1:57(1.06)	2:00(2.33)	✓	✓							M1.1
2017 Aug 21	2:30(1.59)	2:32(3.12)	✓	✓							B7.3
2017 Aug 22	4:00(3.45)	4:04(5.54)	✓	✓							B6.2
	8:31(8.01)	8:37(9.09)	✓	✓							B5.5
2017 Aug 23	6:24(5.48)	6:27(7.27)	✓	✓							C1.6
2017 Aug 24	4:29(3.59)	4:33(5.03)	✓	✓							B1.8
2017 Aug 25	7:24(6.45)	7:28(8.16)	✓	✓							C5.5
2017 Sep 4	1:11	1:12	✓	✓							level-C
	2:11(1.88)	2:32(2.59)	✓	✓							C1.1
2017 Sep 5	4:43(4.03)	4:47(5.37)	✓	✓							M3.2
	6:40(6.03)	7:08(7.13)	✓	✓							M3.8
2017 Sep 6	7:31(6.59)	7:39(8.18)	✓	✓							C2.7
	8:49(8.27)	9:10(9.71)	✓	✓							X2.2
2017 Sep 7	2:00(1.31)	2:03(2.33)	✓	✓							C1.8
	2:52(2.25)	2:54(3.01)	✓	✓							C2.3
	5:32(4.29)	5:38(5.38)	✓	✓							M2.4
	6:17(5.59)	7:11(7.12)	✓	✓							C8.2
	7:25(7.03)	7:49(8.06)	✓	✓							C3.7
	8:22(8.02)	8:39(9.07)	✓	✓							C1.7
	9:16(8.40)	9:28(9.54)	✓	✓							C2.3
2017 Sep 8	1:25(1.06)	1:47(2.12)	✓	✓							C5.3
	2:11(1.49)	2:10(2.05)	✓	✓							C1.2
	2:40(1.49)	2:49(2.59)	✓	✓							M1.3
	3:27(3.00)	3:34(4.07)	✓	✓							C1.4
	3:39(3.09)	3:47(4.15)	✓	✓							level-C
	4:20	4:26	✓	✓							C1.4
	5:08(5.01)	6:57(7.11)	✓	✓							C8.3 C2.9 C1.7
	7:03(6.34)	7:11(7.40)	✓	✓							C6.0
	7:20(7.10)	9:33(9.28)	✓	✓							M8.1
	9:36(9.06)	9:39(10.11)	✓	✓							C1.3
2017 Sep 9	5:14	5:19	✓	✓							M1.1 in decay phase
	6:52(6.21)	6:57(7.28)	✓	✓							C1.7

Year	2013	2014	2015	2016	2017	2018	2019	2020	Total (Tb)
Data Size(Tb)	0.14	4.4	12.2	58	152	121	~40	~5	~390

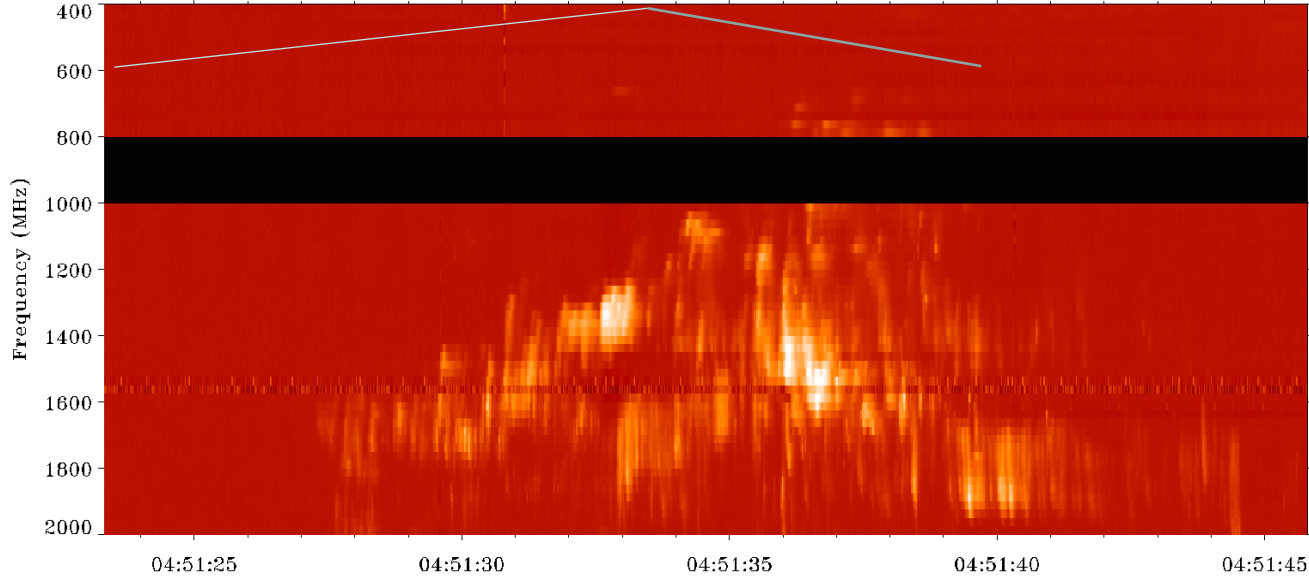
Zhang, Ming-Hui, et al, Observational results of MUSER during 2014–2019 , 2021, RAA , Vol. 21 No. 11, 284(12pp)

The 20151122 event

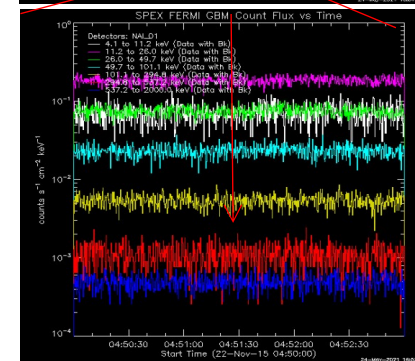
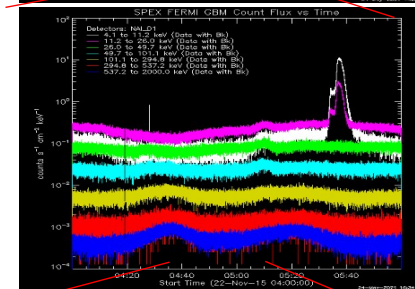
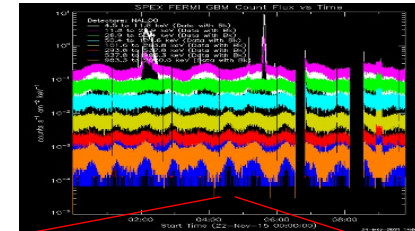
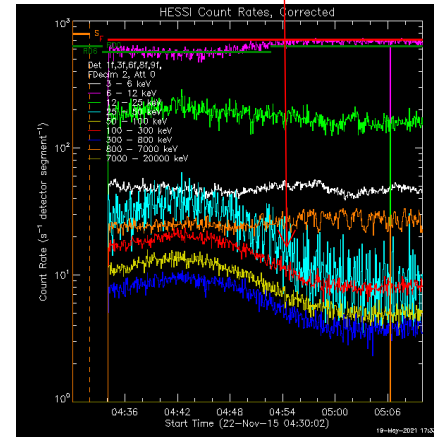
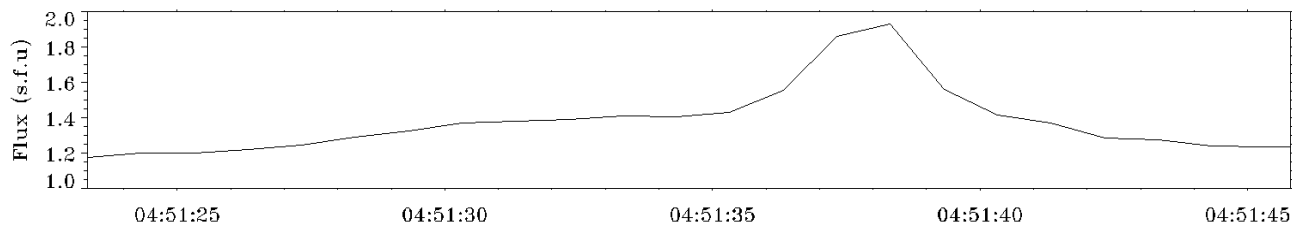
RHESSI and Fermi HXR showed no increasing



The cross correlation Spectrum of all baselines (after flag) on 20151122



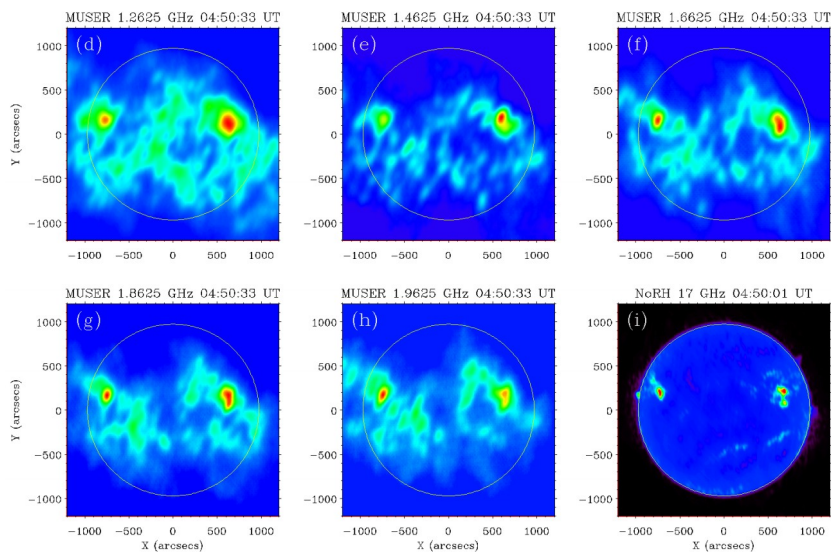
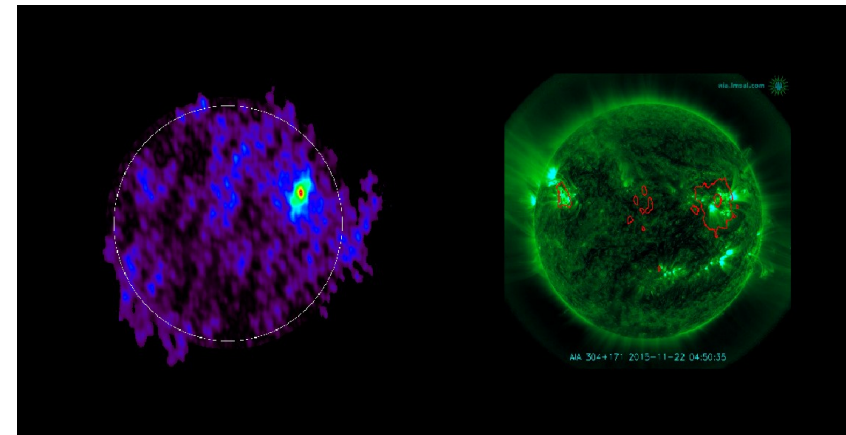
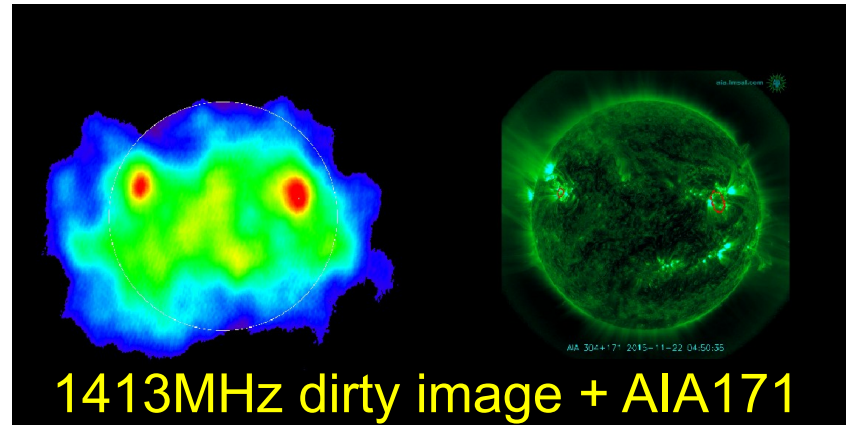
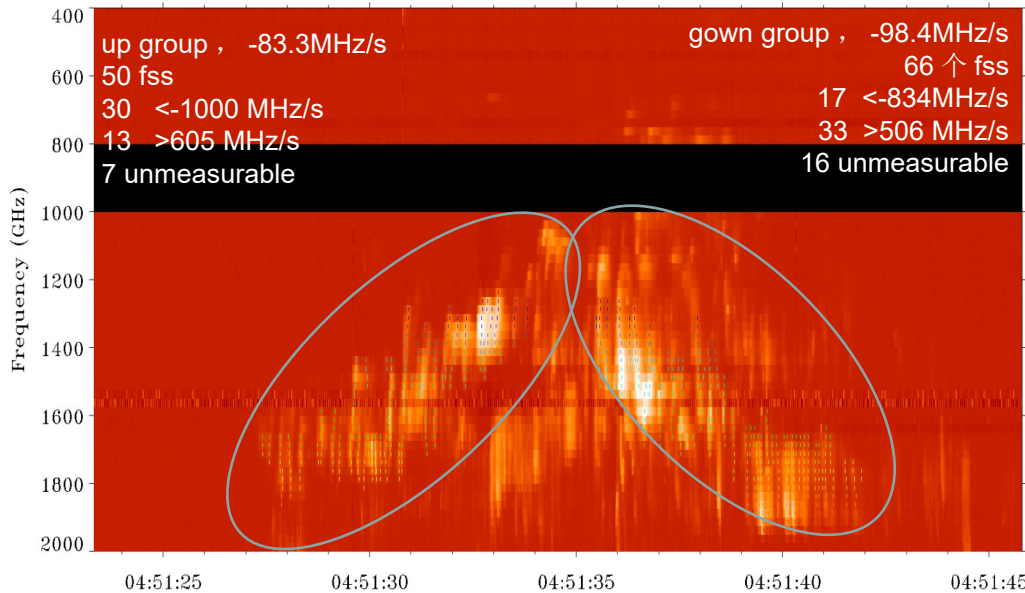
Norh 17GHz flux profile at active region on 20151122



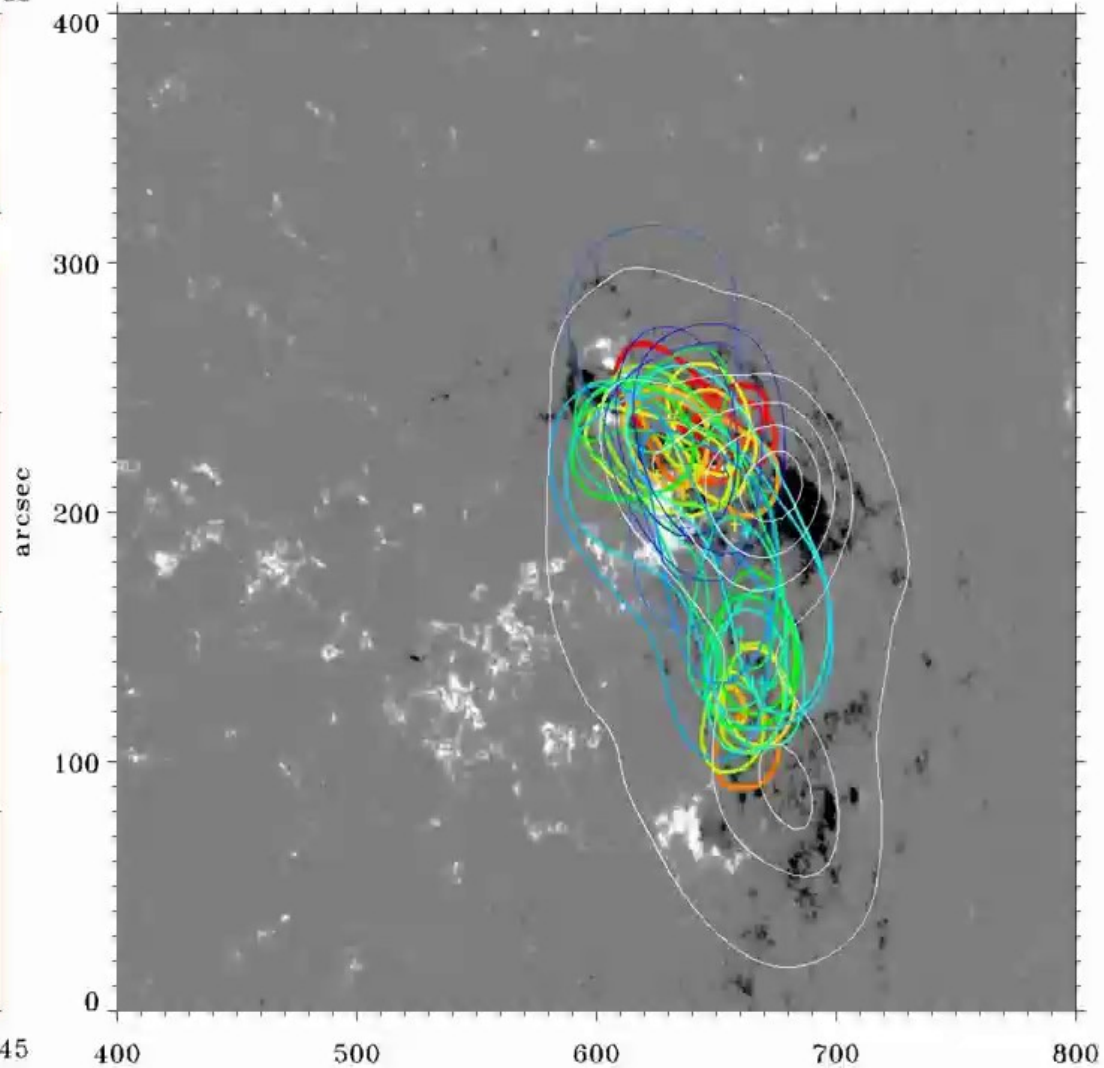
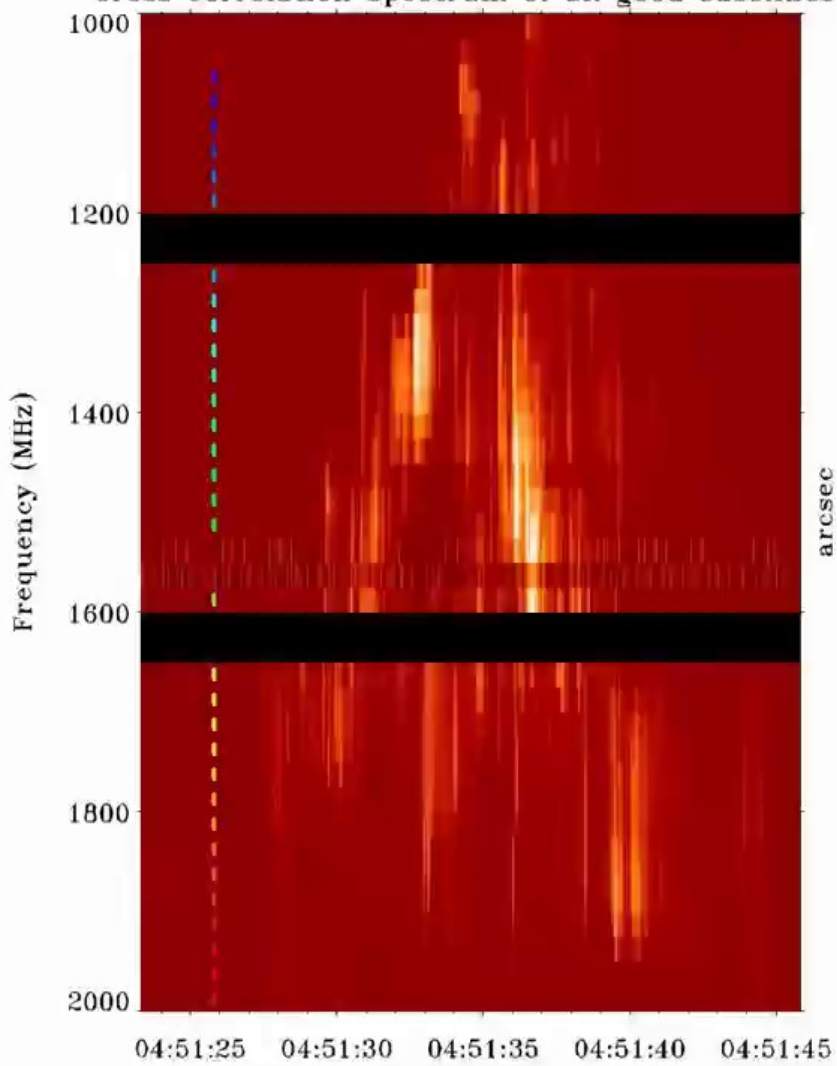
spectrum and image

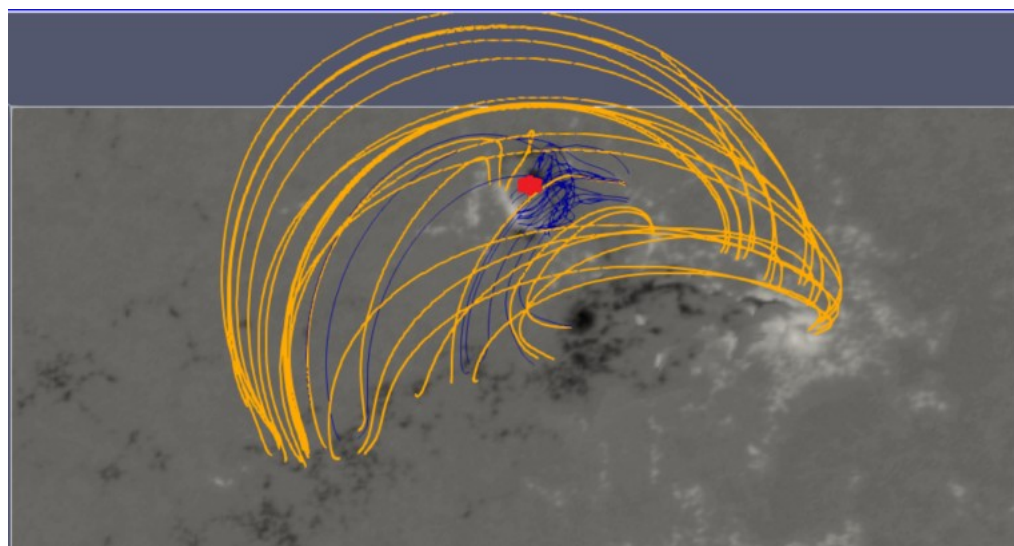
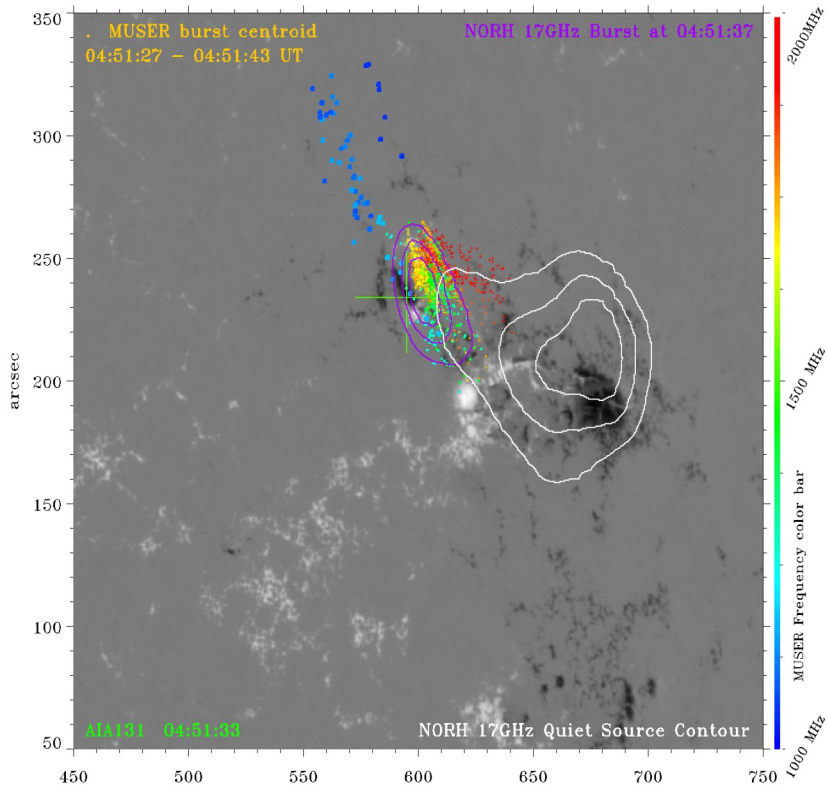
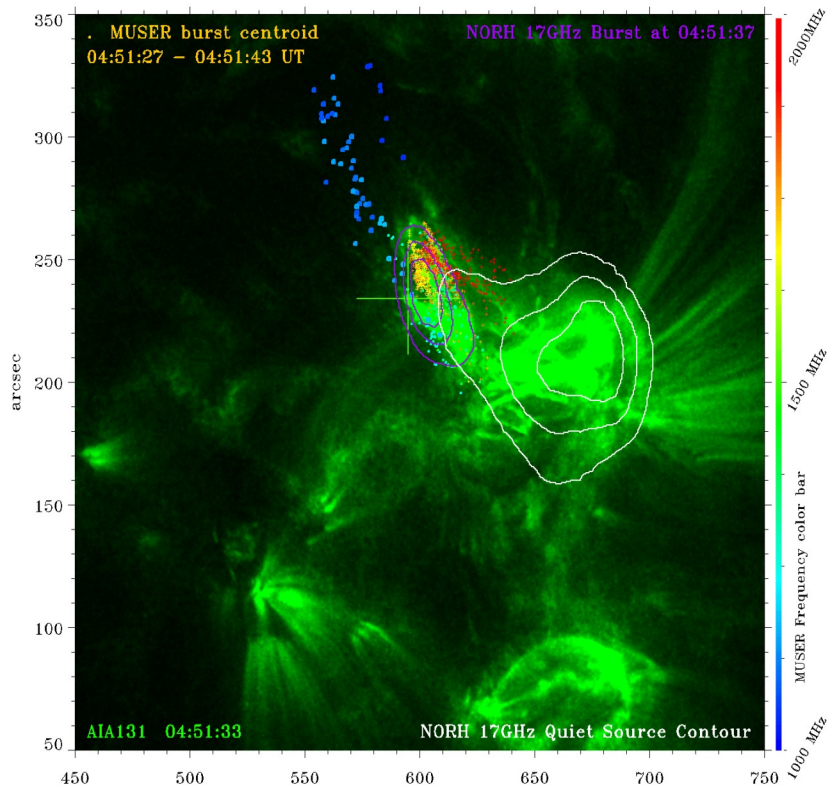
drifting

The cross correlation spectrum of all baselines with bursts on 20151122



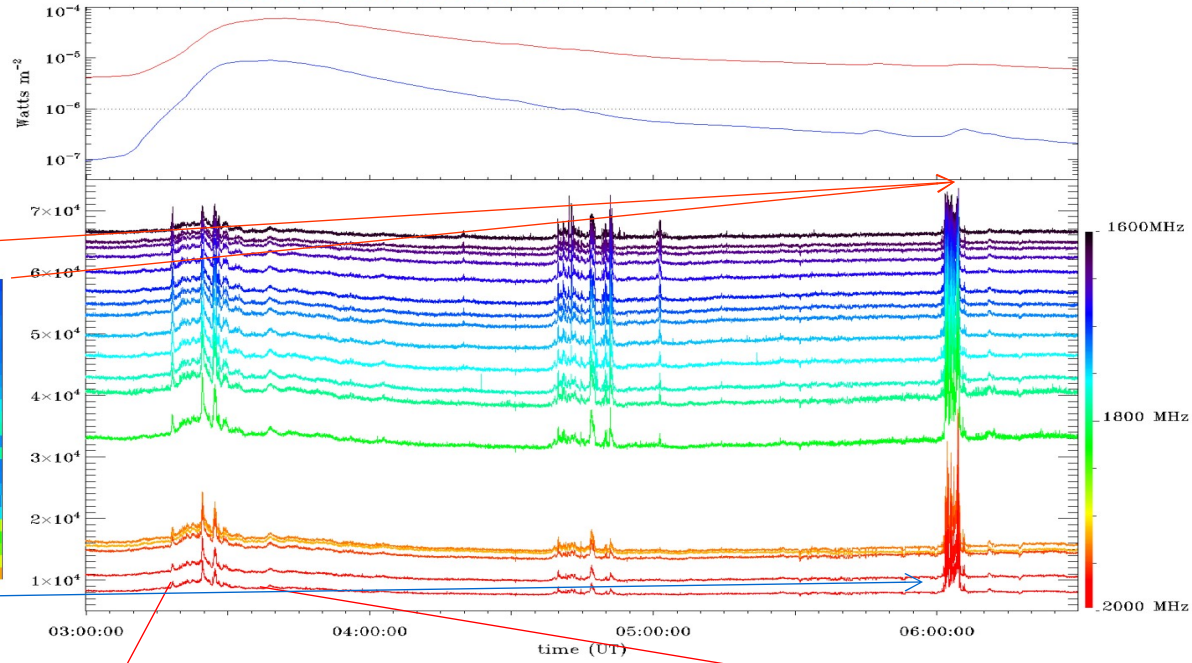
Cross Correlation Spectrum of all good baselines



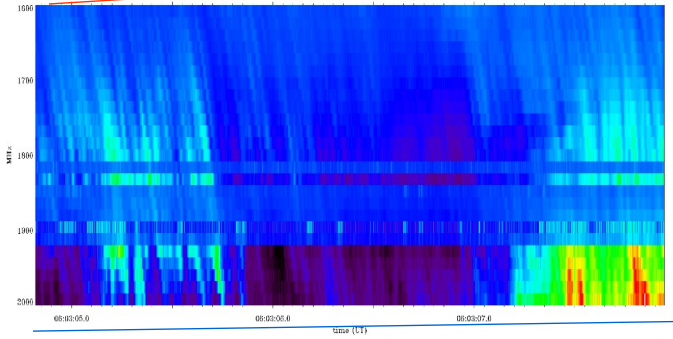


MUSER-I upgrade

GOES flux and MUSER Observation on 2023-01-15 Antenna (IA1)

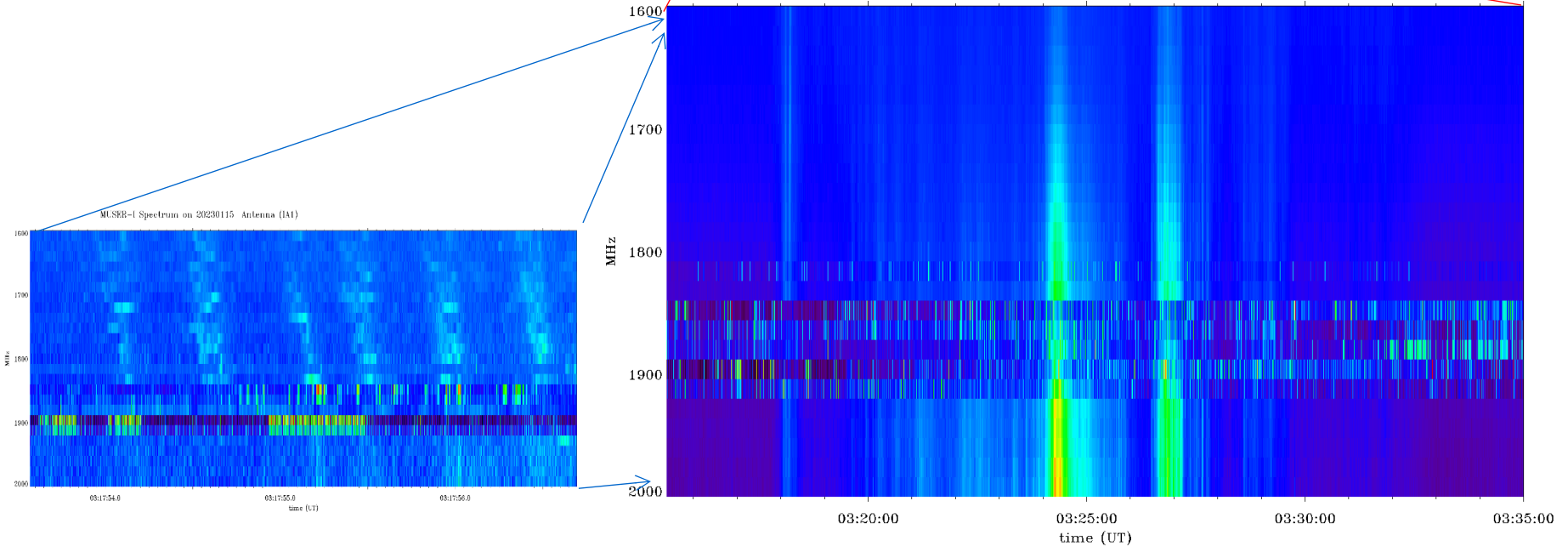


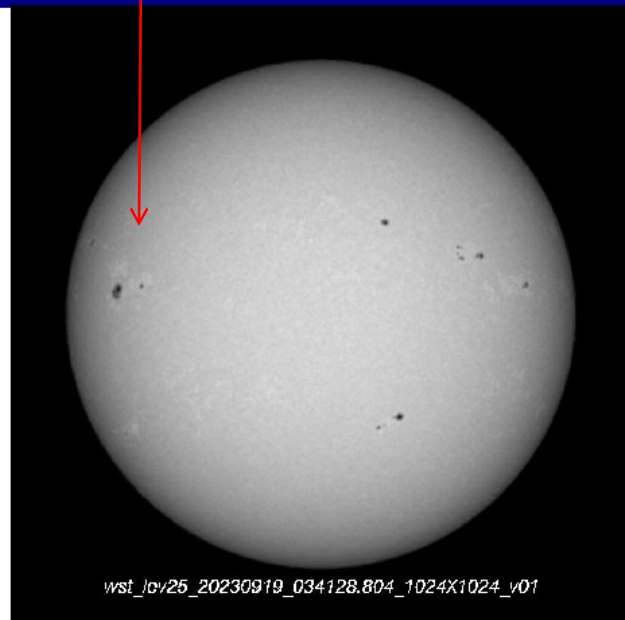
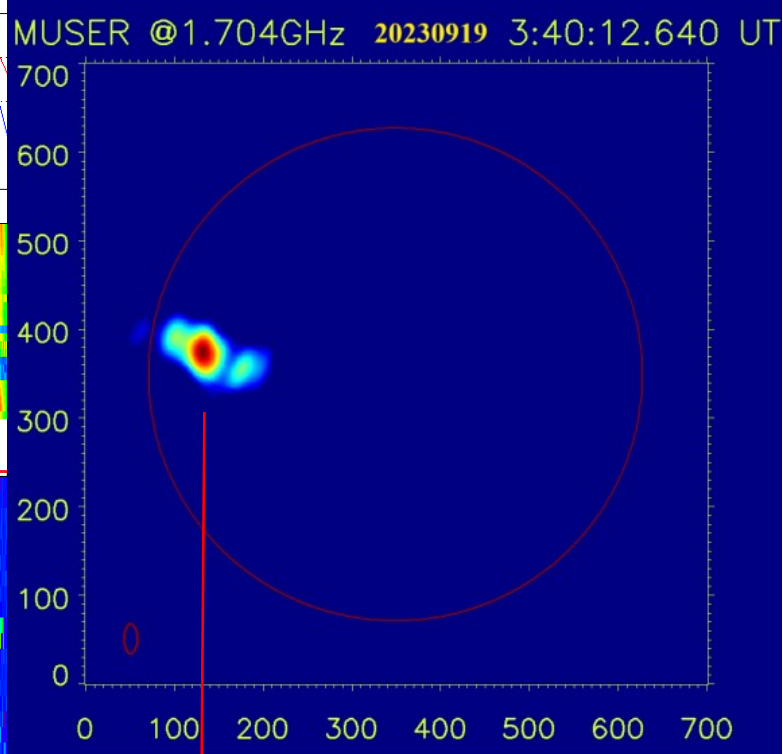
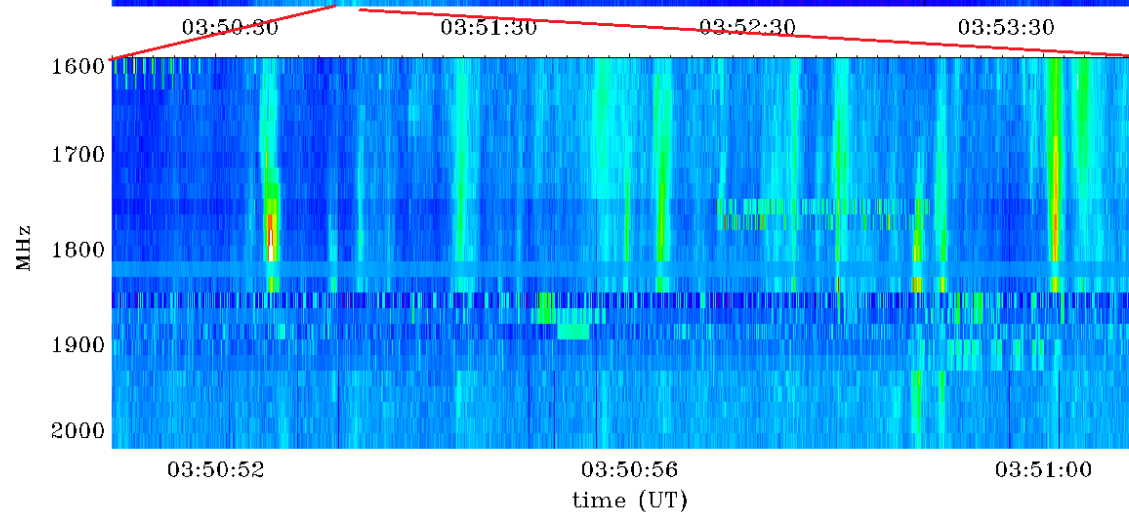
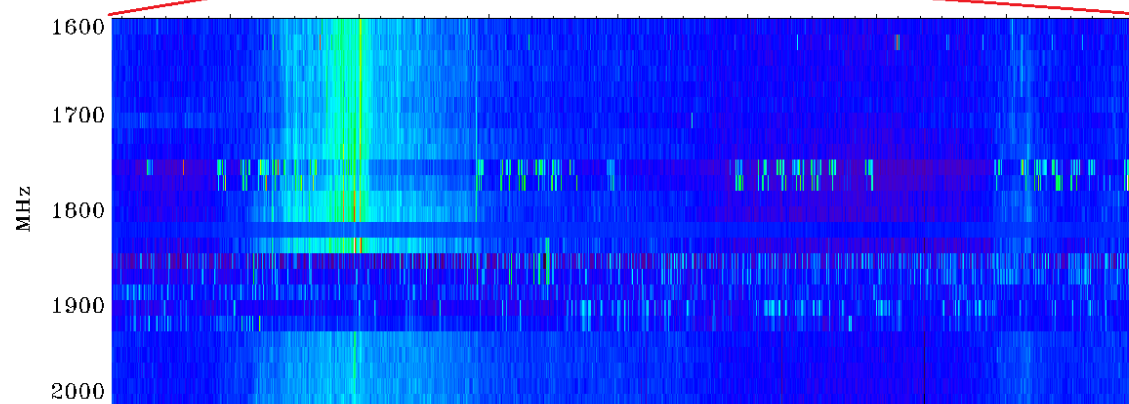
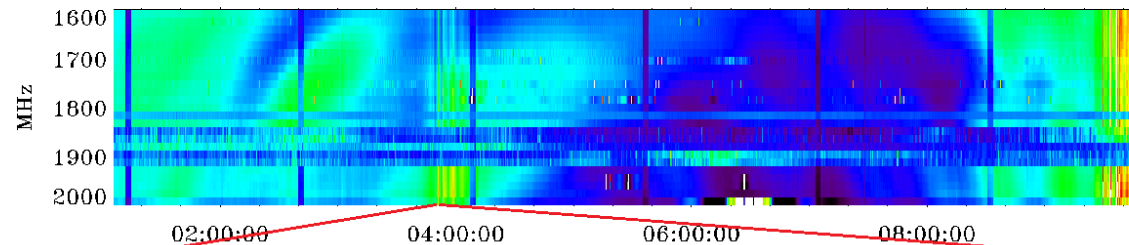
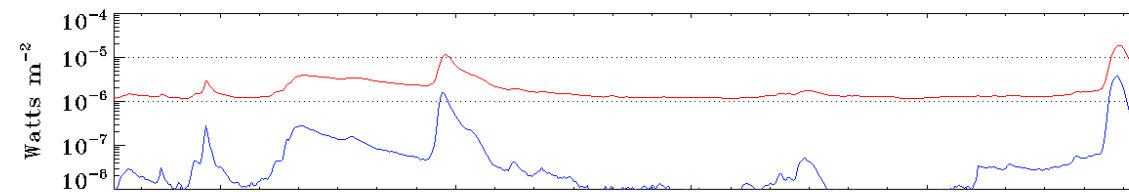
MUSER-I Spectrum on 20230115 Antenna (IA1)



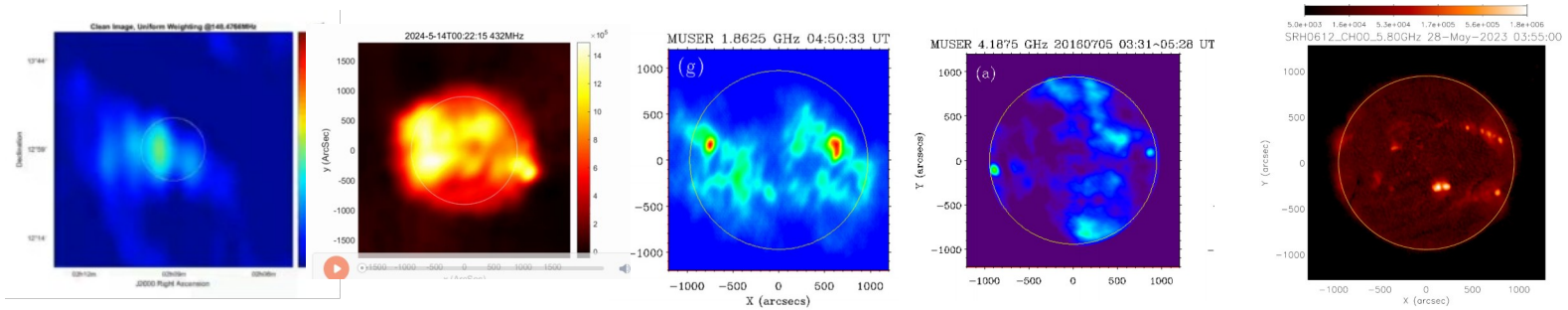
fiber burst

MUSER-I Spectrum on 20230115 Antenna (IA1)





Joint study with SRH and CBS



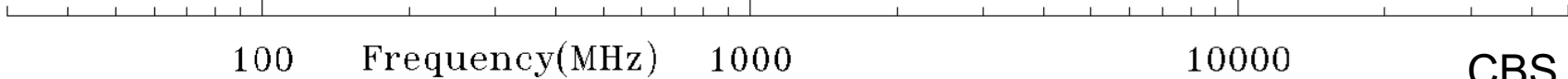
DSRT

SRH-L SRH-M SRH-H

MUSER-L

MUSER-I

MUSER-H (TBU)



CBS

CBS



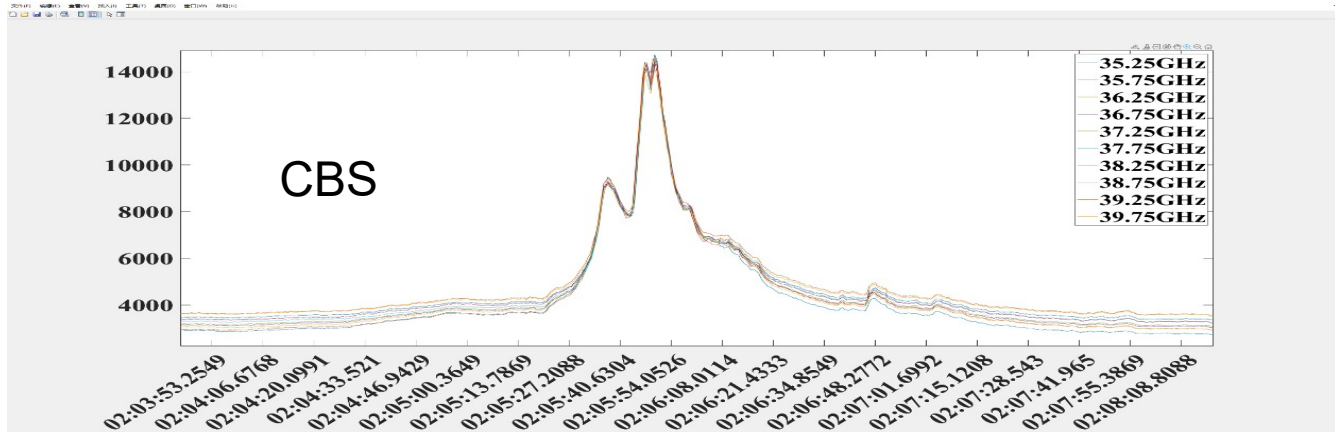
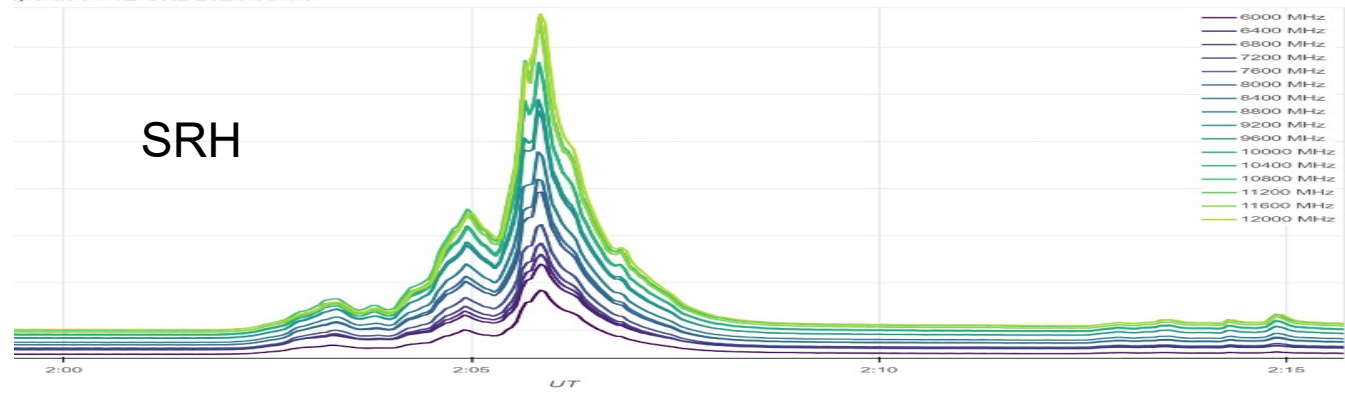
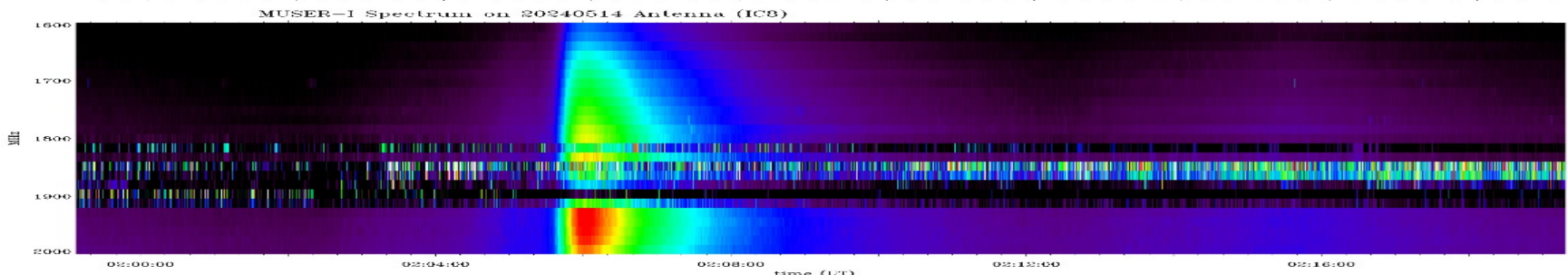
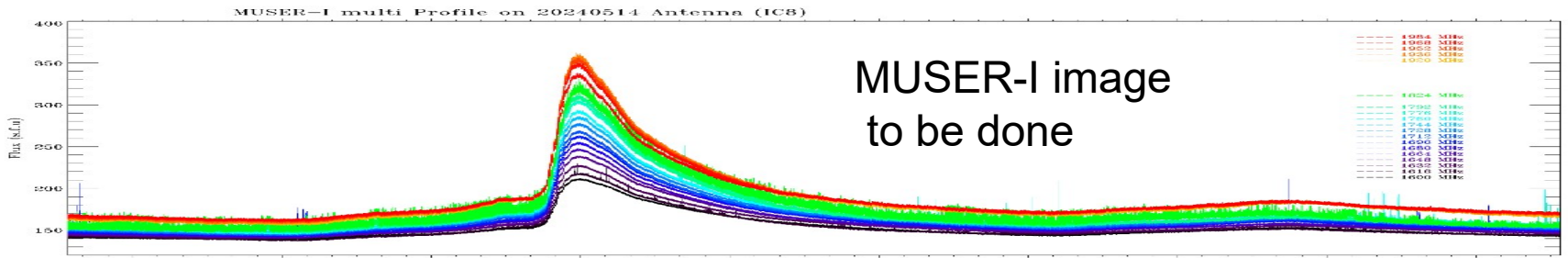
event list

20230111 , 0140-0220UT,
 20230112, 0335-0350UT
 20230115, 0310-0400UT , 0430-0500UT , 0550-0610UT
 20230208, 0251-0254UT
 20230209 , 0245-0300UT (weak) , 0715-0717UT (ffs)
 20230210 , 0240-0310UT , 0730-0830UT

.....

SRH event	Time	GOES	Location	Other data	Notes	MUSER-I	MUSER-L
2024-01-23	03:30	M5.1	[-270, 576]	KW	short, classical light curve	yes	no
2024-01-23	07:15	C9.6	[600, -213]		long rise phase	yes	no
2024-01-29	04:25	M6.8	[883, 422]	KW	~1 h, classical	yes	yes
2024-02-06	03:00	M4.3	[720, -560]		very long and complicated event (up to 3 peaks)	yes	yes
2024-02-07	03:40	M5.2	[730, -576]	DSRT, CSO	~1 h, classical	yes	no
2024-02-09	07:31	C9.4	[-228, -150]		very short, classical, late time	yes	yes
2024-02-10	03:40	M3.4	[-50, -100]	DSRT, CSO	~30 m, classical + secondary event	yes	no
2024-02-12	03:45	M6.5	[422, -115]		short, classical	yes	yes
2024-02-16	06:55	X2.5	[890, -340]	KW	short, classical; late time	yes	yes
2024-02-22	06:40	X1.7	[570, 350]	KW	~30 m, classical	yes	yes
2024-02-24	06:31	M4.5	[-190, 422]		short, classical, late time	yes	yes
2024-03-17	05:40	C3.8	[-883, 345]		~30 m, classical (+ a sympathetic weaker flare)	yes	yes
2024-03-20	07:40	M7.4	[-720, -192]		~1 h, classical + secondary event, late time	yes	yes
2024-03-23	01:30	X1.1	[-192, -115]		long and complicated event, early time	no	no
2024-03-23	07:10	M2.5	[-192, -115]		short, classical + secondary event	yes	yes
2024-03-24	06:05	M2.2	[38, -115]		very short, classical	yes	yes
2024-03-25	06:45	M4.4	[120, -120]	KW	extremely short, classical	yes	yes
2024-03-26	00:35	M1.8	[500, -110]	KW	extremely short, early time	yes	no
2024-03-28	06:30	M7.1	[730, -192]		~30 m, classical	yes	?
2024-03-29	02:20	M3.2	[810, -190]		short, classical	yes	yes
2024-04-01	01:30	M3.9	[850, 255]		very short + long secondary	yes	no
2024-05-14	02:09	X1.7	[S17W89]			yes	no





Summary

- MUSER had recorded a huge amount of data (>4PB). The event list and figure would be organized. It will promote the joint study with SRH.
- We have started the data analysis for one candidate event. The joint analysis will be done in detail for several common view solar flares, and determining the radio flux, peak frequency, spectral or spatial characteristic. It will trace the source and process of energy release and particle acceleration.
- The spectral and spatial analysis will be helpful in the study of Solar-Terrestrial space weather in wide radio spectral range of 30MHz - 40GHz.



Спасибо !

