



SINP MSU FEASIBILITIES FOR IMCP

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OUTLINE

International Meridian Circle Program (IMCP) tasks
where SINP MSU can contribute:

- Space Weather (SW)
- Global Electric Circuit (GE)
- Geomagnetic Field Variations (GM)



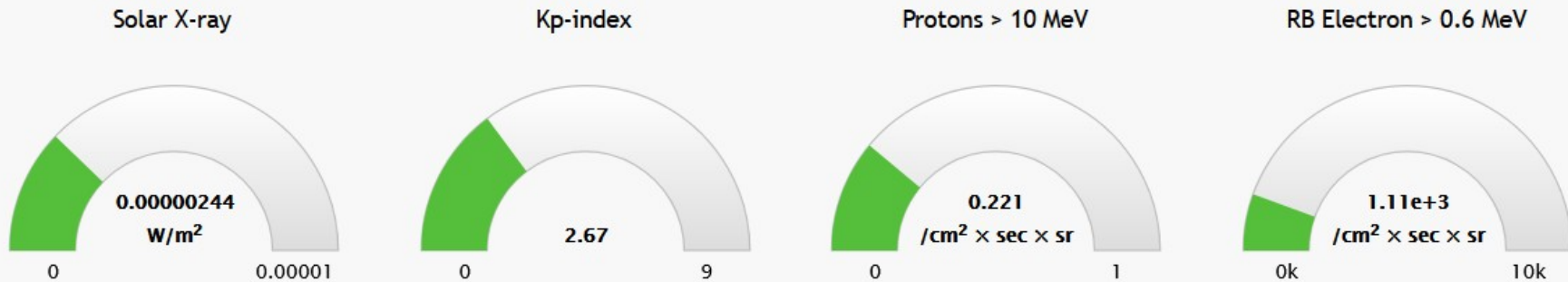
SWX <https://swx.sinp.msu.ru/index.php>



SINP MSU Space Weather Analysis Center

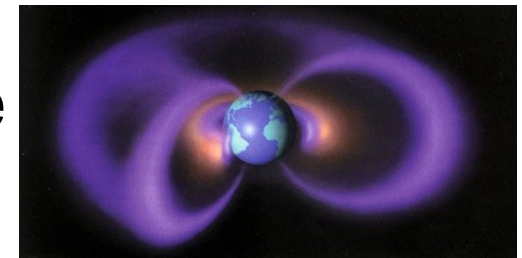
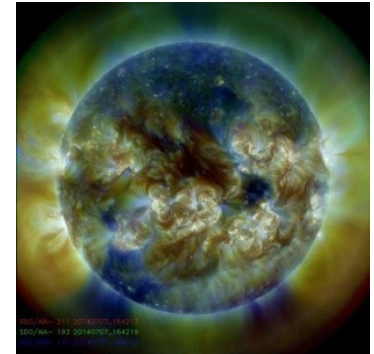
Space Weather Analysis Centre of SINP MSU provides information about the current state of near-Earth's space. Information Services ([SWX](#)) on the website of the center provide access to current data describing the level of solar activity, geomagnetic and radiation state of the magnetosphere and the heliosphere in real time. For data analysis, the models of the space environment, working in off-line as well as on-line mode have been implemented. Interactive services allow one to retrieve and analyze data in a given time moment. [SWX](#) is a flexible system for the analysis and forecasting of space weather in the near-Earth's space.

Current conditions in space (2024-08-27 17:16:58 UT)

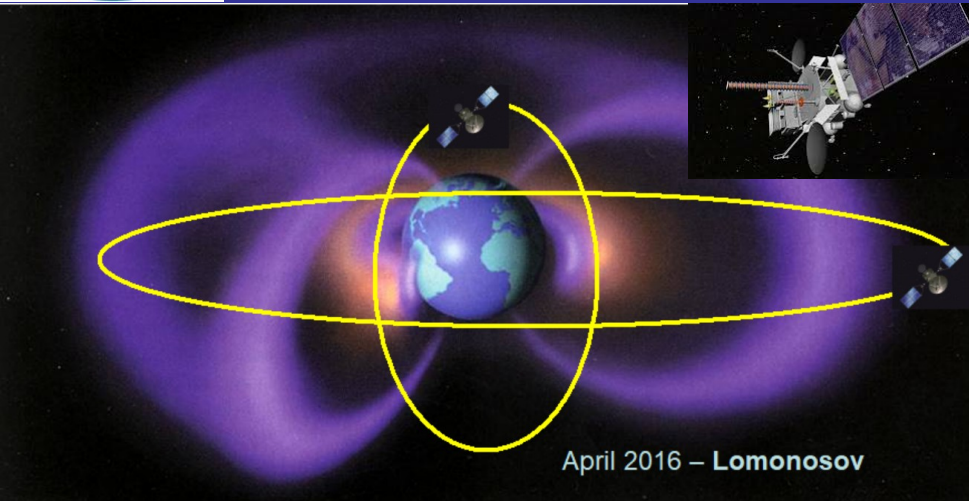


Space weather diagnostic

- Sun observations
- Solar wind parameters
- Particle fluxes in the magnetosphere



Operational services: Data + Models



GEO: Electro L1 & L2

76°E

SKIF-6, SKL-E (SINP MSU)

SEP Protons 1 - 320 MeV

ERB Electrons 0,03 - 20 MeV



HEO: Arctica M1 & M2

40,000km 160/340E

MSGI-M, SKL-M (SINP MSU)

SEP & ERB Protons 2 - >160 MeV

Electrons: 0,15 - 10 MeV

Plasma 0.15 - 20 keV



LEO: Meteor-M1, M2, M2-4

Sun-synchronous 3/15, 4/16, 9/21 MLT

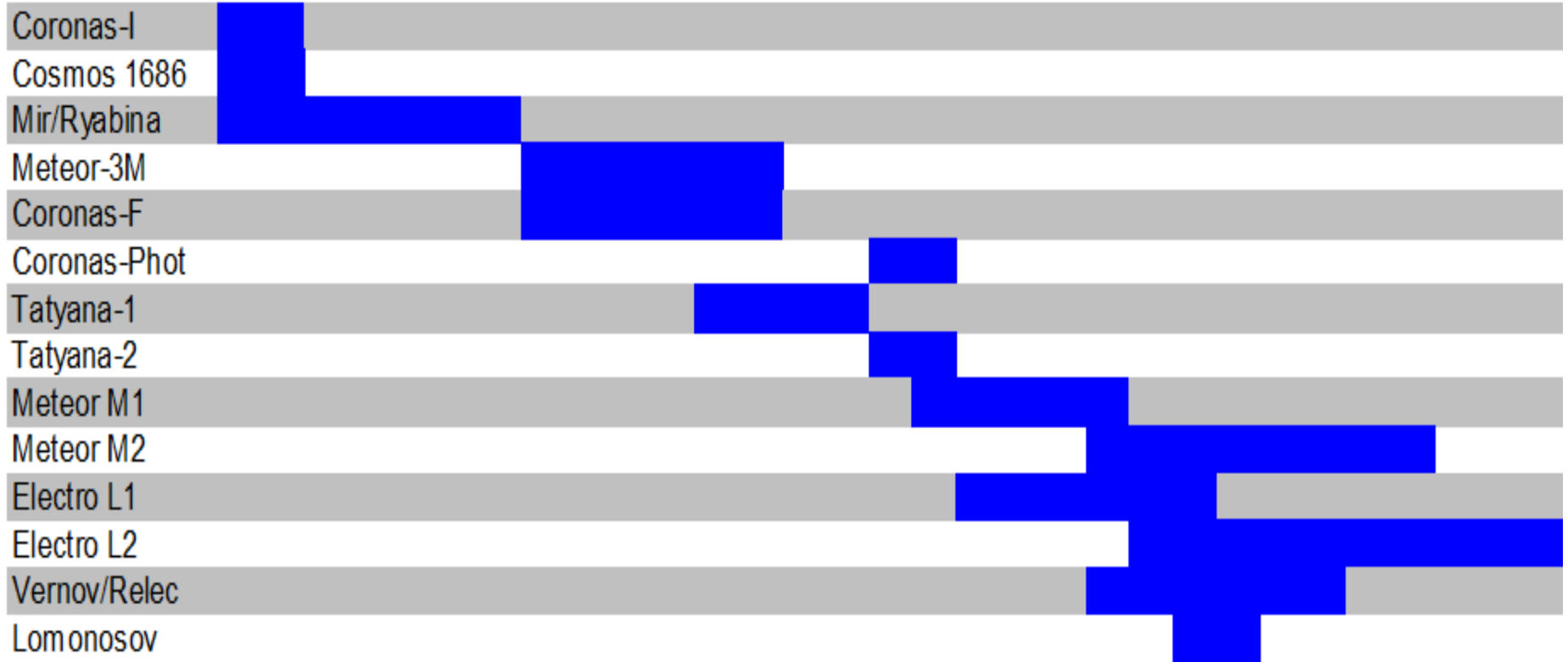
MSGI-M, SKL-M (SINP MSU)

SEP & ERB Protons 1 - 160 MeV

Electrons: 0,03 - 13 MeV (auroral 0.03-16 keV)

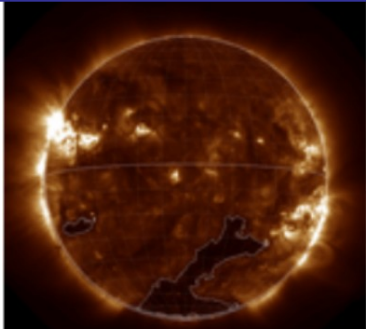
Missions and Data availability

94 95 96 97 98 99 00 01 02 03 04 05 06 07 08 09 10 11 12 13 14 15 16 17 18 19 20 21 22 23 24



Currently operating sc are used but their data are still not public:

- Electro L4
- Arktika M1 & M2
- Meteor M24

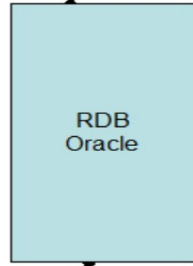


Coronal holes

Solar wind

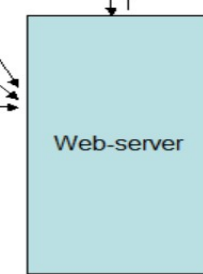
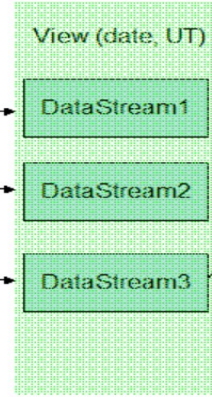
SW velocity (middle-term)

Meteor M1
Electro L1,
Meteor M2



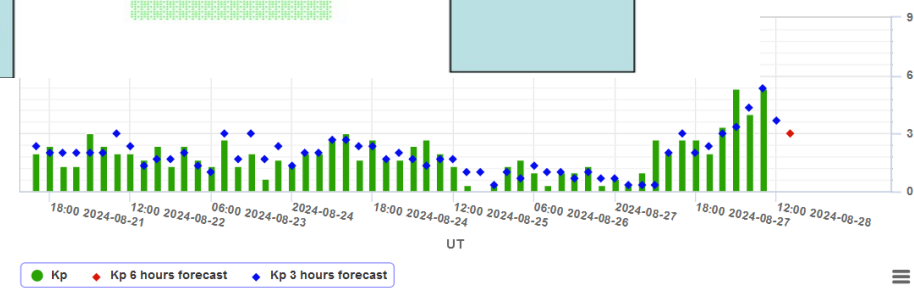
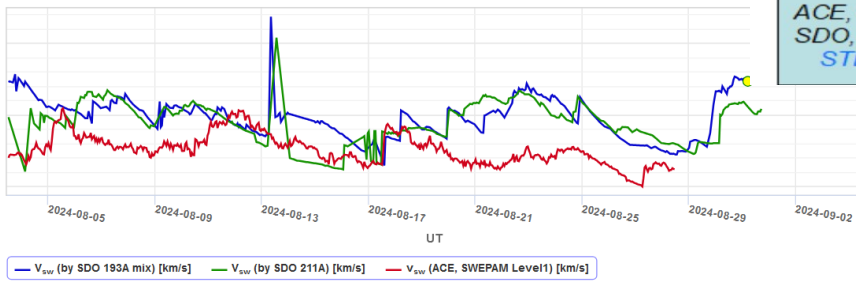
ACE, GOES,
SDO, indices,
STEREO

Models



Kp

client



Dst forecast

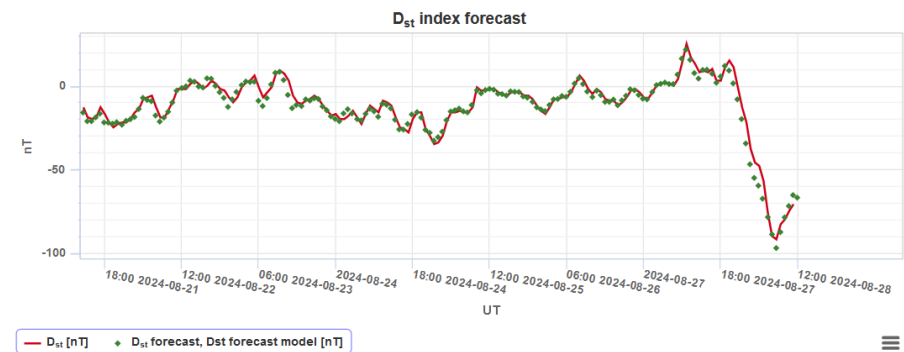
Dst

Geosynchronous Relativistic electrons daily fluxes

Меню
Дата
2015/12/14 11:21
submit



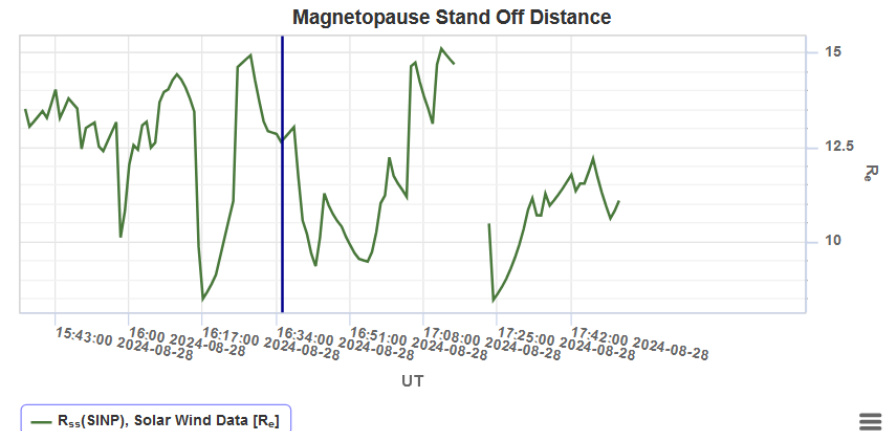
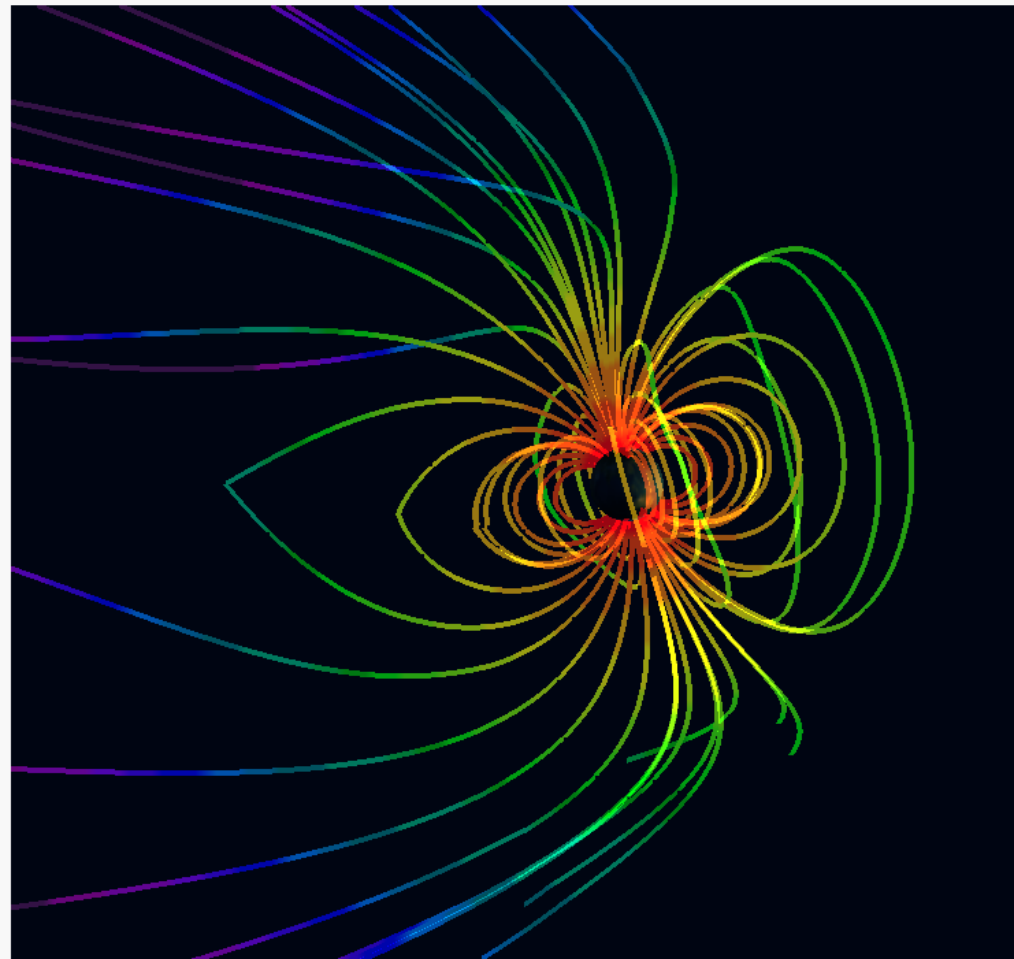
2024-08-21 15:55:23 - 2024-08-29 14:55:23



3D-magnetosphere

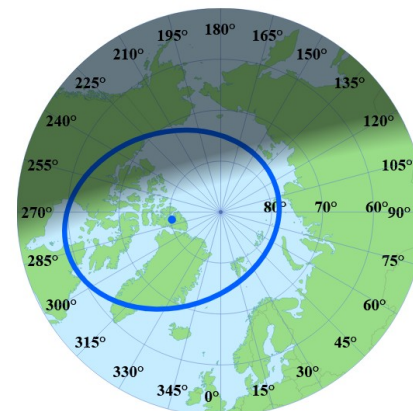
<http://swx.sinp.msu.ru/3d.php?lang=en>

Magnetopause standoff distance

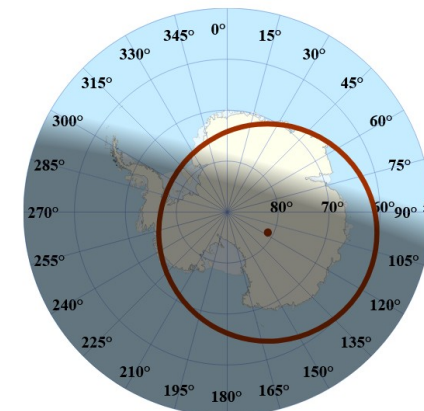


ORB high-latitude boundary

Northern hemisphere



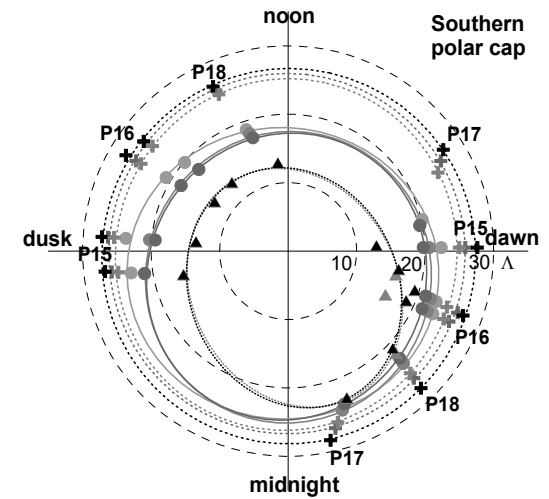
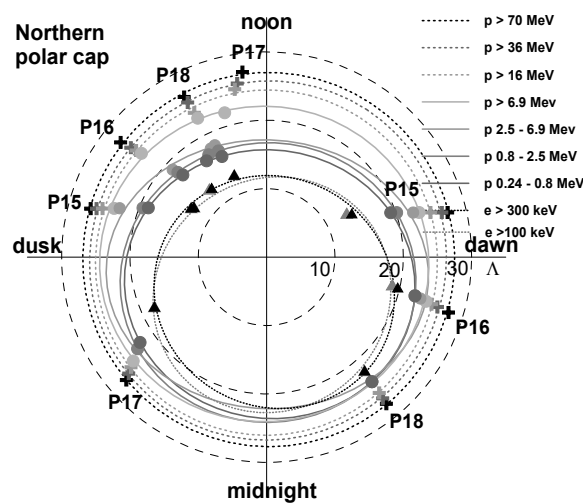
Southern hemisphere



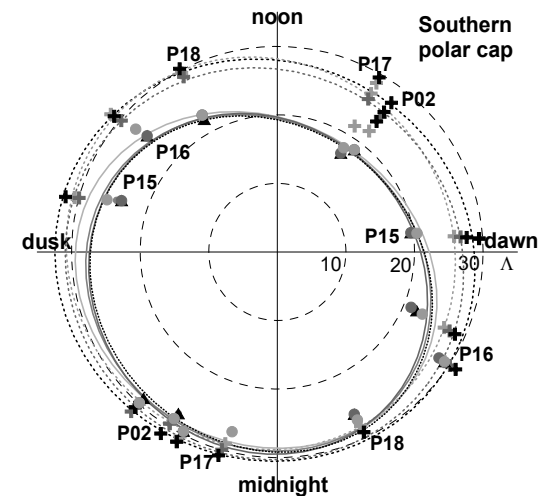
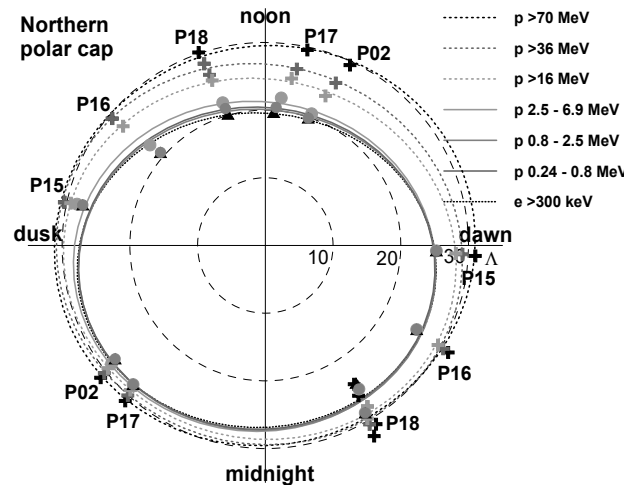
Enable Earth Rotation Mode

Invariant co-latitudes of SEP cutoff

Quiet time



Magnetic storm



Approximation of the SCR cutoff boundaries under quiet conditions and during the magnetic storm at 17:00 UT on December 14, 2006 in the northern (left) and southern (right) hemispheres. The cutoff latitudes determined for HEP, LEP, and electrons are indicated by crosses, circles, and triangles, respectively. The flybys of various spacecraft are indicated by their respective abbreviations.

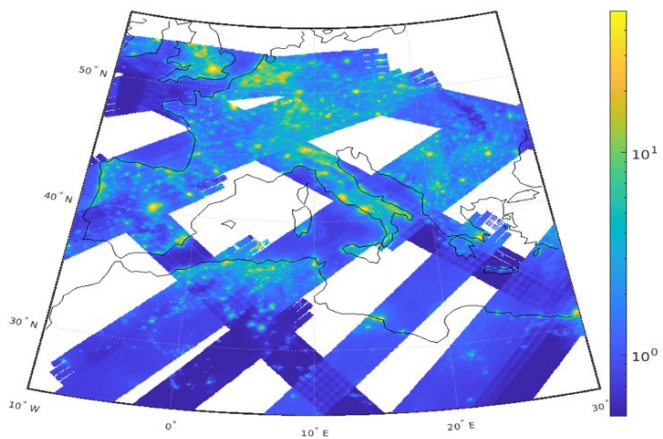
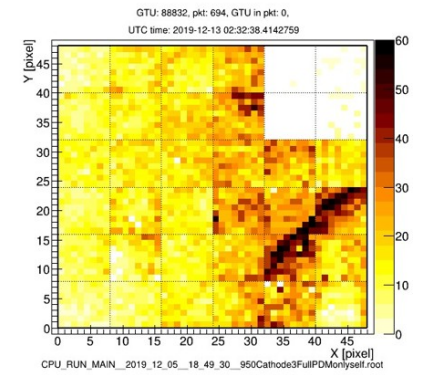
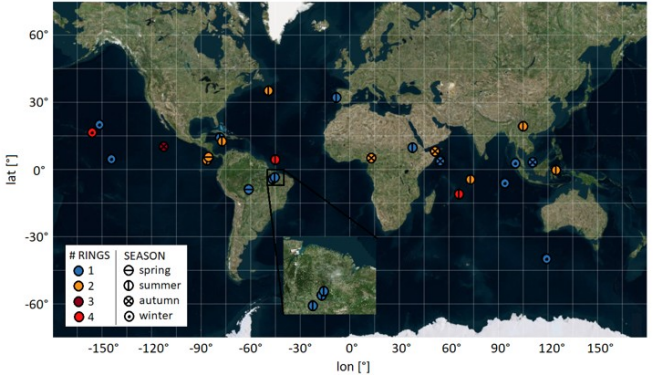
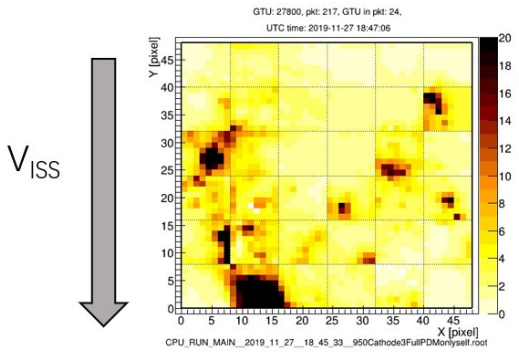
Dimensions	370×370×620 mm
FOV	36°× 36° = 0.42 sr
Entrance pupil diameter	25 cm
Focal distance	25–27 cm
Pixel size	3 mm
Number of pixels	2304
Spatial resolution at ground level	6 км
Area of observations	260 км × 260 км = 6.8·10 ⁴ км ²
Temporal resolution	2.5 us, 320 us, 40 ms



uf-atmosphere.html
 Launched on 22.08.2019, delivered to the ISS on 27.08.2019.
 To date, 110 experimental sessions have been conducted, the equipment operates well.
 Multi-level trigger system with different time resolution allows to measure variable atmospheric phenomena.

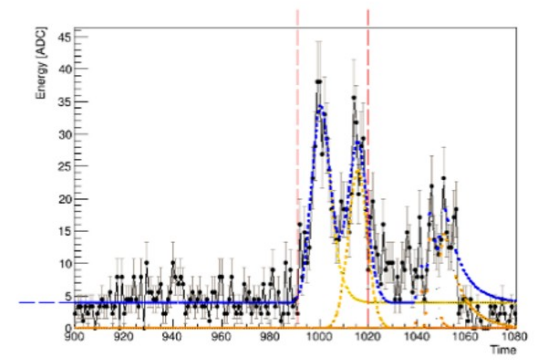
Monitoring of the UV emission and Transient luminous events

ELVES map and one example of fine spatio-temporal structure



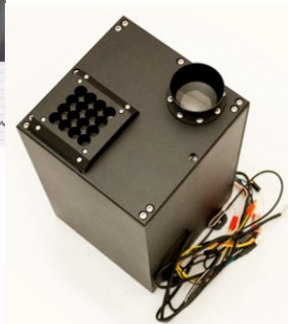
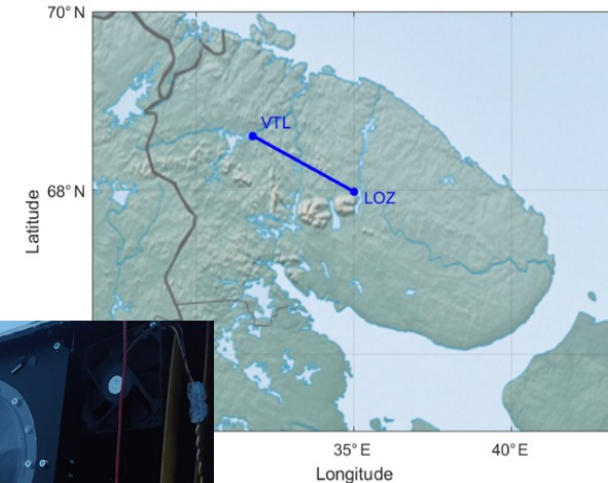
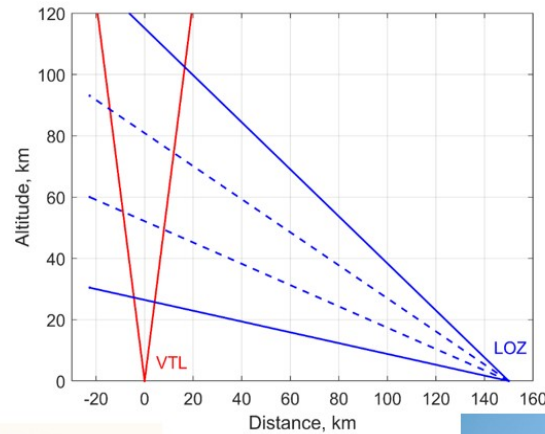
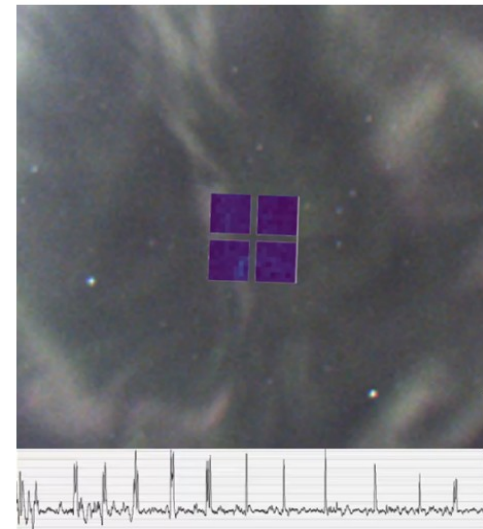
ELVE light curve

Unique measurements of the fine spatiotemporal structure of ELVES, allowing probing intra-cloud processes.



PAIPS (Pulsating Aurora Imaging Photometers System)

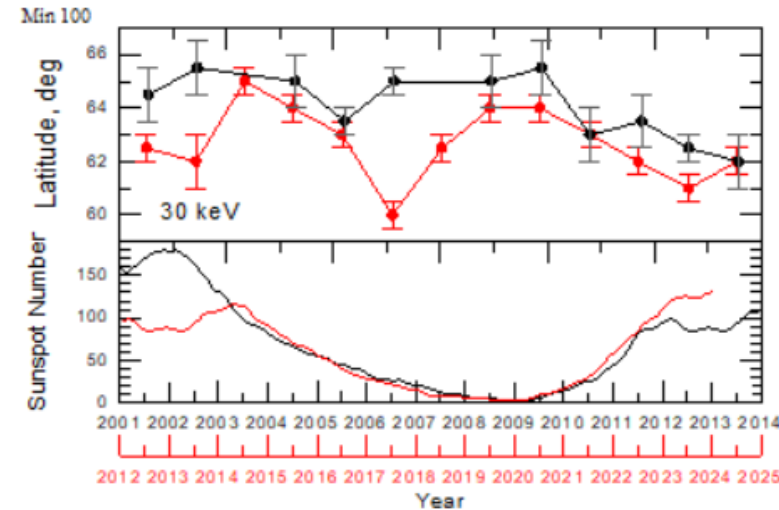
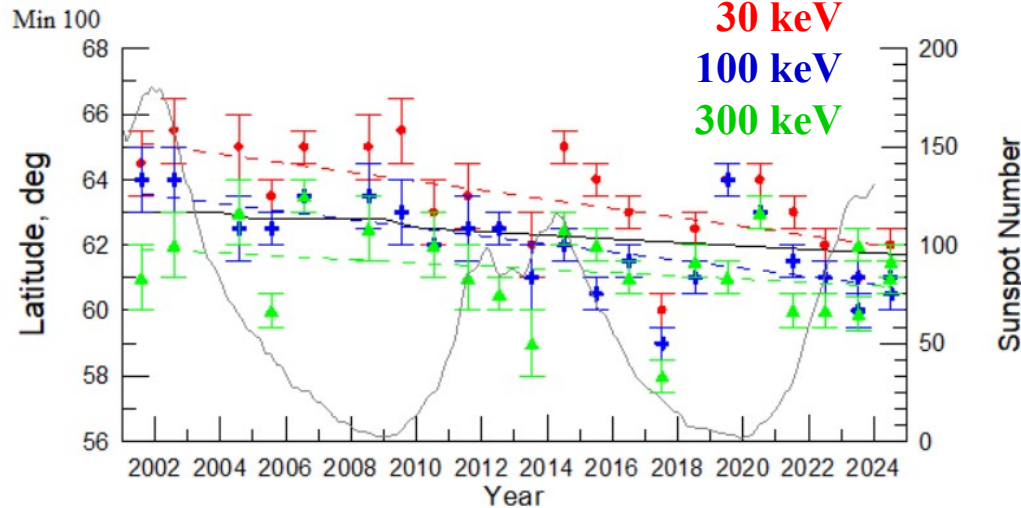
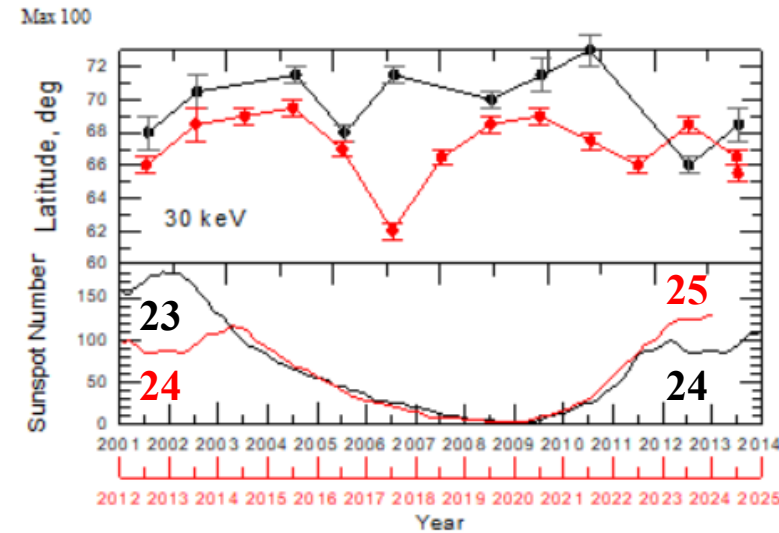
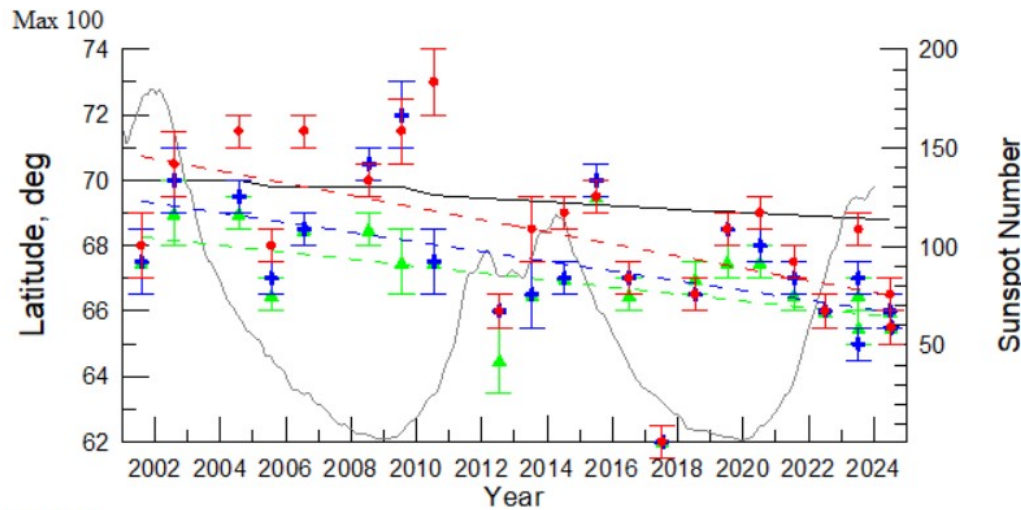
<https://uhedr.sinp.msu.ru/paips.html>



Two highly sensitive photometers at a distance of 150 km at the Verkhnetulomsky and Lovozero observatories, which allow stereometric measurements of atmospheric pulsations and UV-microbursts with high time resolution, complementing the information from all-sky cameras about the fine spatiotemporal structure of the emission and spectrum of precipitating particles.

Tracing the geomagnetic field by the location of ORB

Decadal dynamic of the outer ERB location at 100°E



Geographic latitude of the maximum (top) and inner edge (bottom) of the outer ERB projection at height of ~ 850 km and around 100°E during magnetic quiet days. Dashed curves show a latitudinal shift predicted by IGRF model of corresponding epochs.

SINP MSU can contribute to the following tasks of the IMCP:

Space weather

- Monitoring of the energetic particles from ERB and SEP (p 1- 320 MeV & e 30 keV – 20 MeV)
- Forecasting of the solar wind velocity, geomagnetic indices Kp & Dst, and relativistic electrons at GEO
- Modeling of the manetosphere, magnetopause, outer ERB, SEP penetration and their radiation effects

Global Electric Circuit

- Transient Luminous Events – spaceborne and ground based observations in UV

Geomagnetic field variations

- Tracing of the ERB projections to the low latitudes